

Properties of selected flaring events observed by SphinX and other instruments within AR 11024

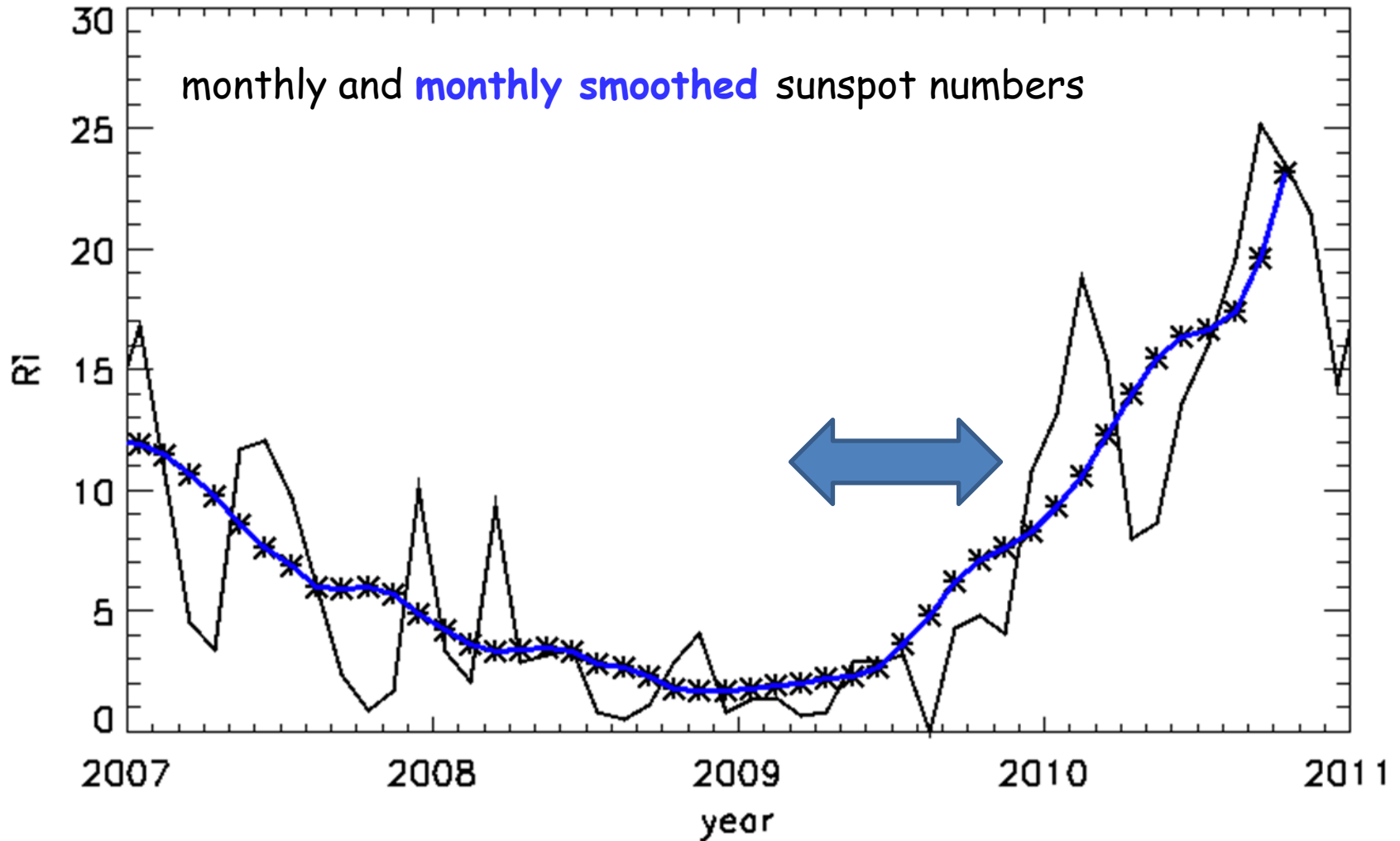
B. Sylwester, J. Sylwester, M. Siarkowski
K.J.H. Phillips, A.J. Engell

*Space Research Center of Polish Academy of Sciences, Wrocław,
Poland*

Mullard Space Science Lab., University College London, UK

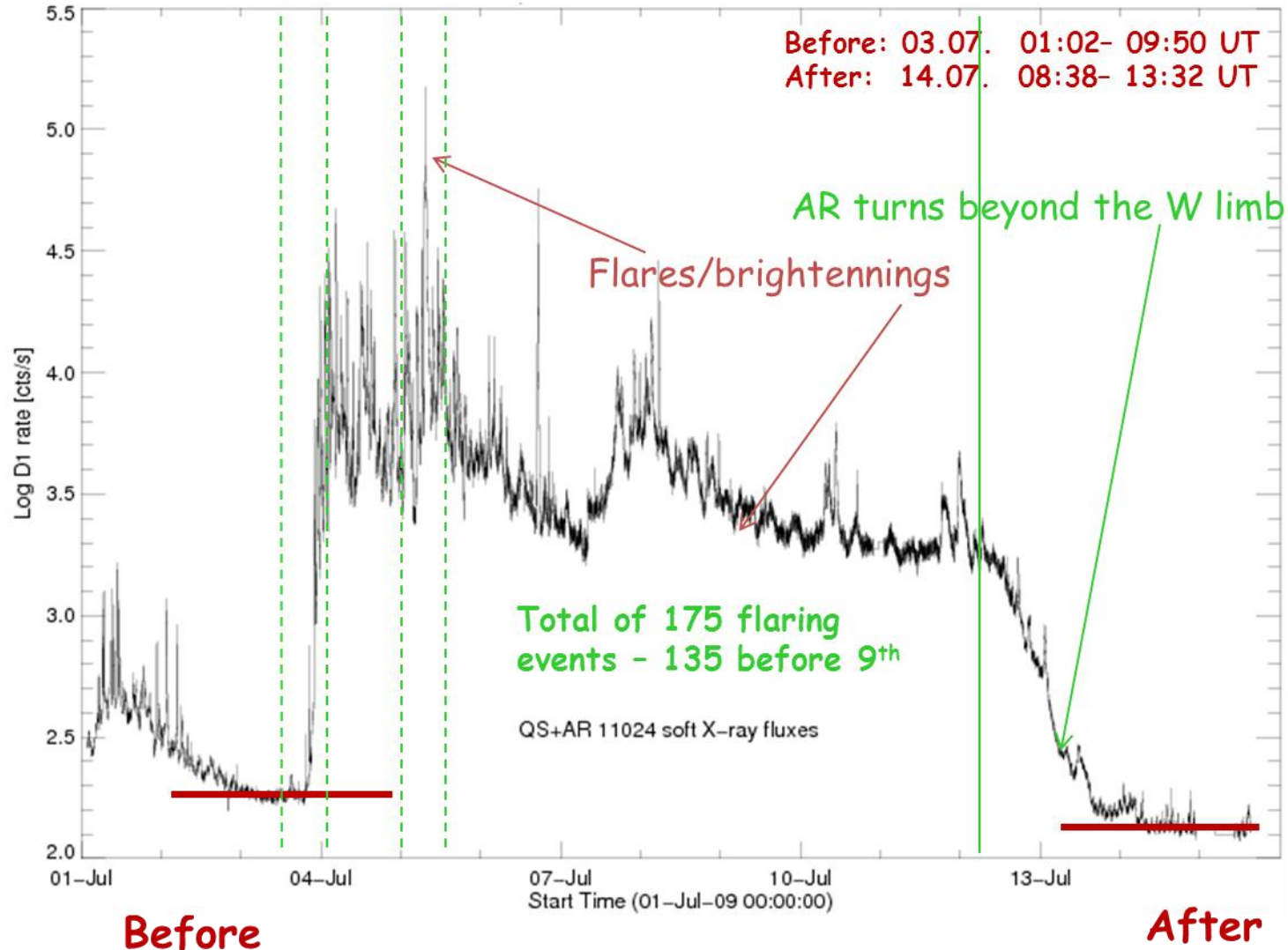
Harvard-Smithsonian Center for Astrophysics, Cambridge, USA

23/24 activity minimum & SphinX

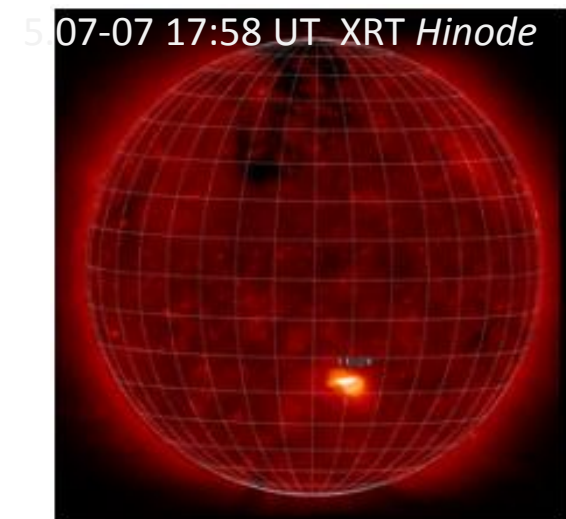
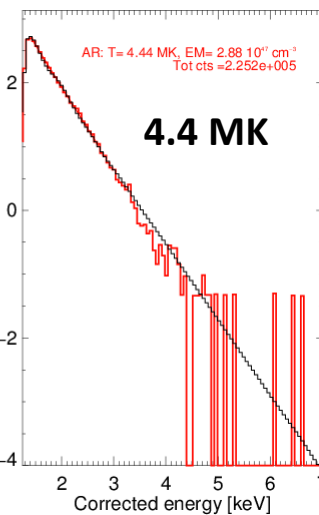
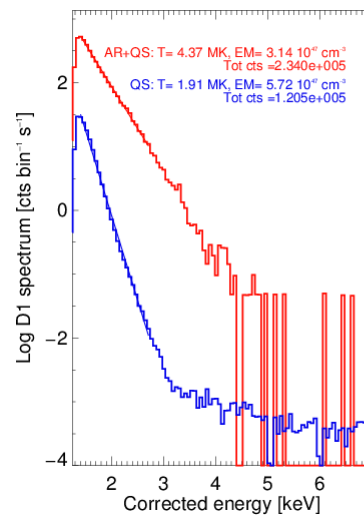
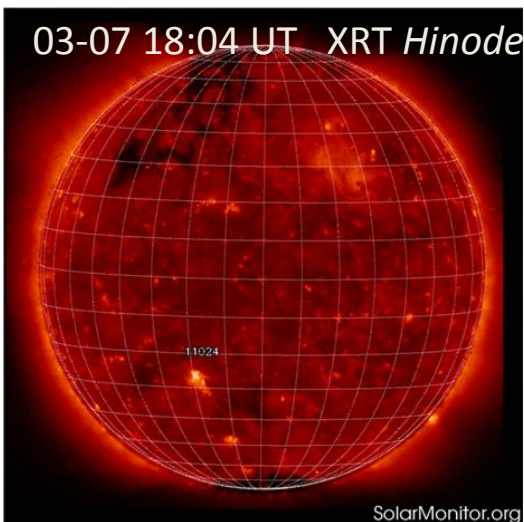
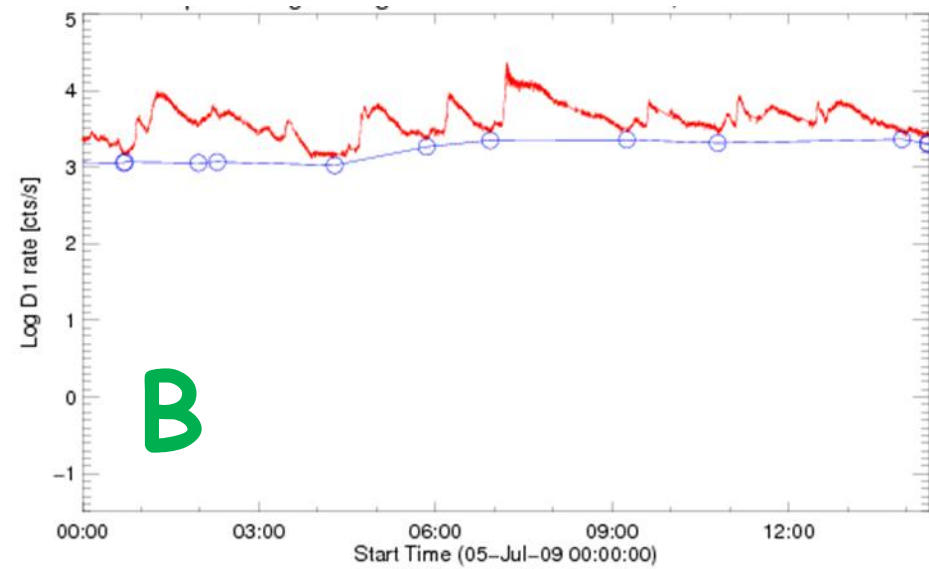
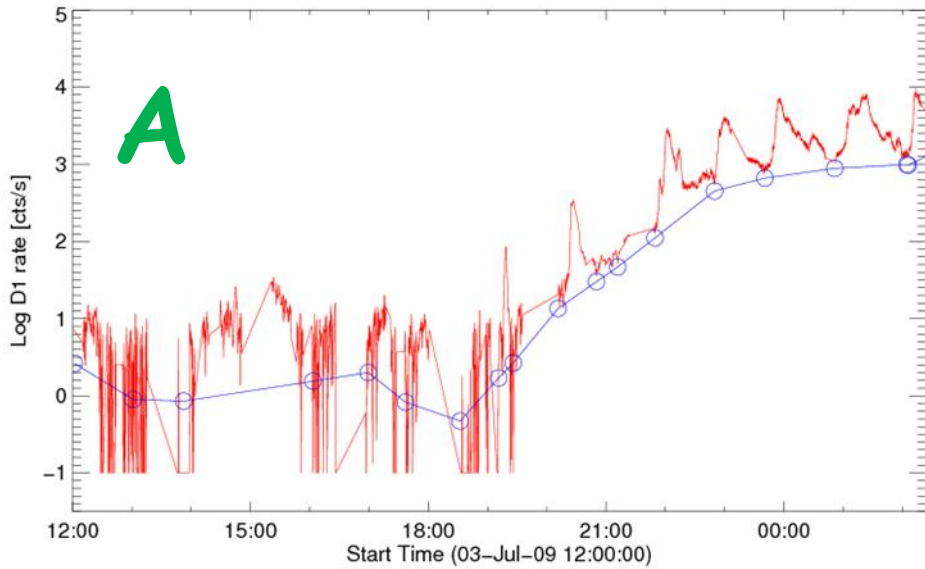


SphinX flux as measured above 1.19 keV (1-16 July 2009)

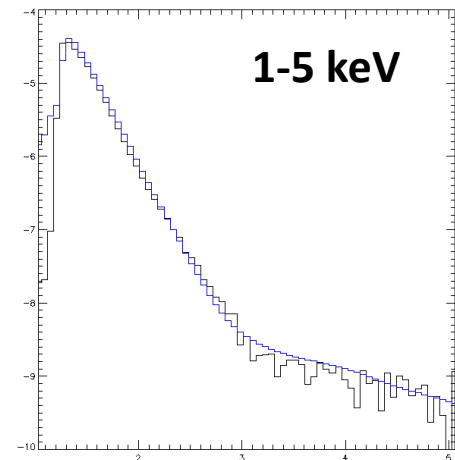
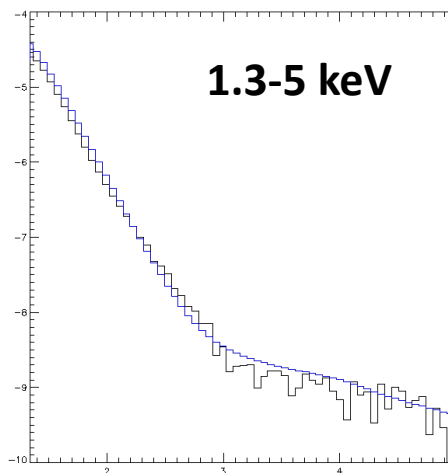
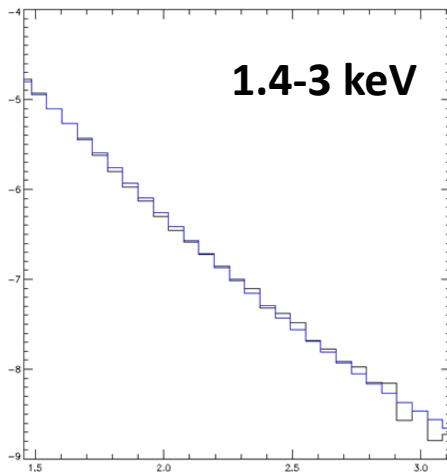
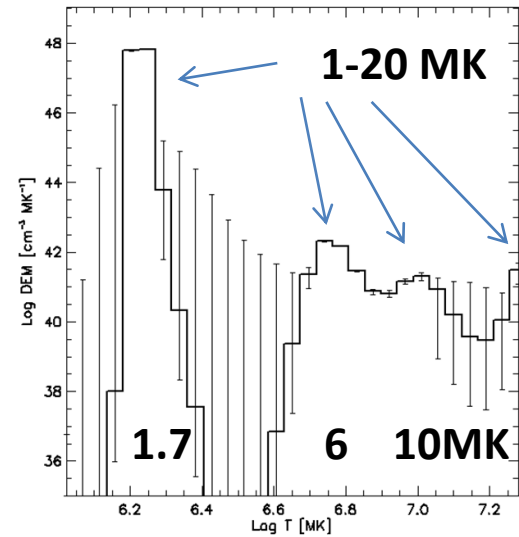
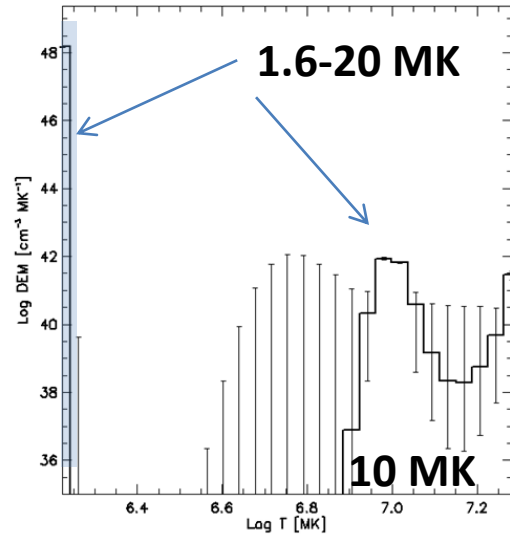
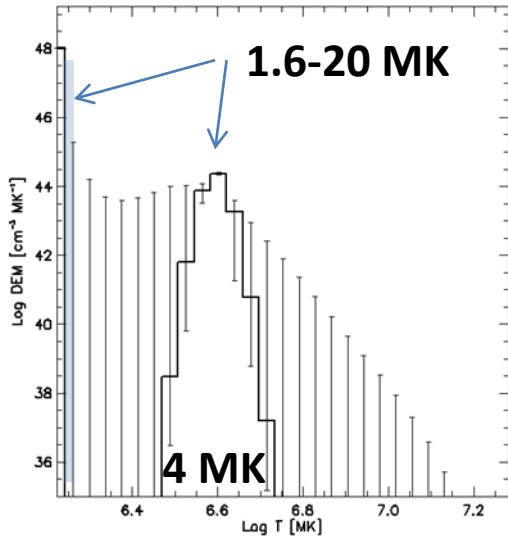
A B



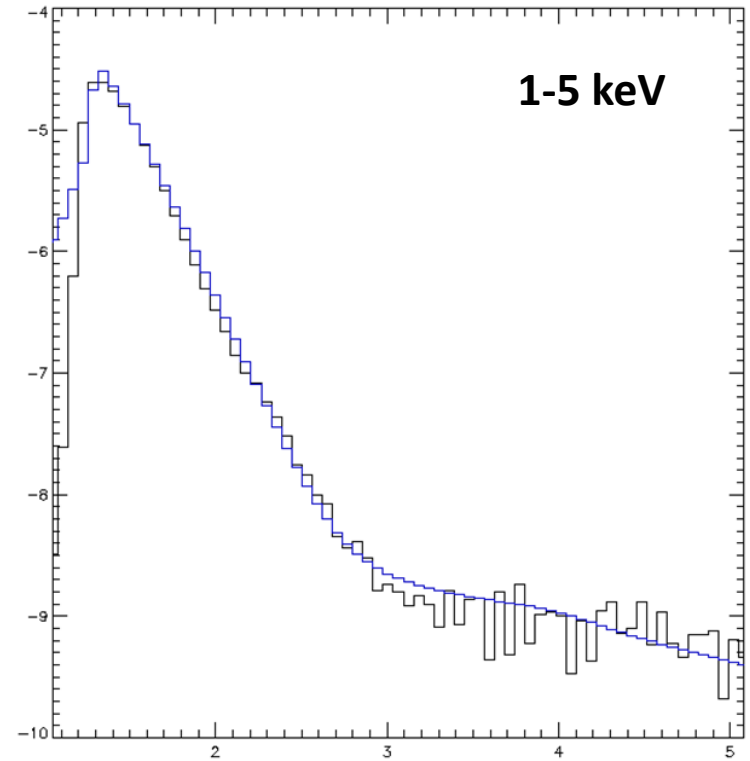
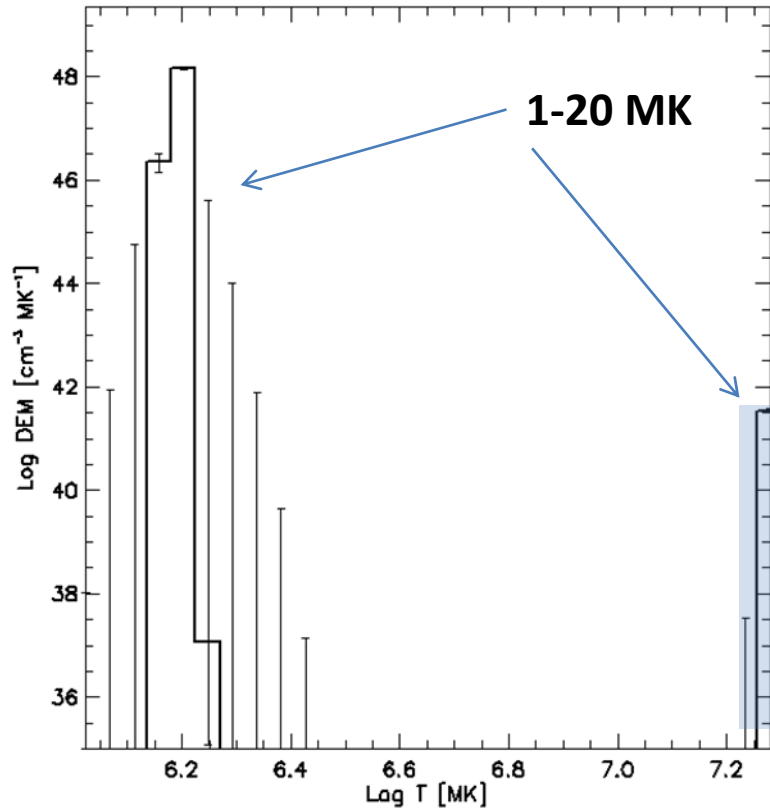
Subtraction of AR contribution (226 points selected)



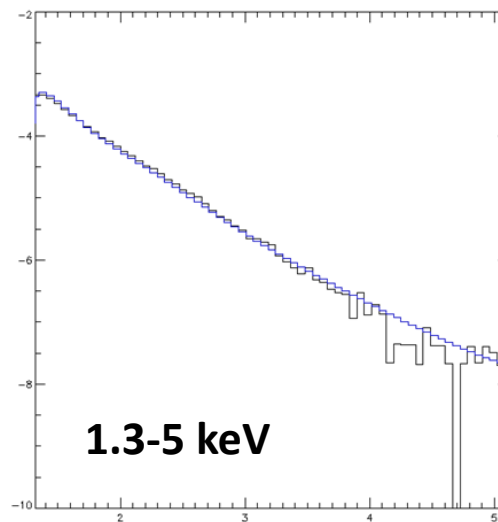
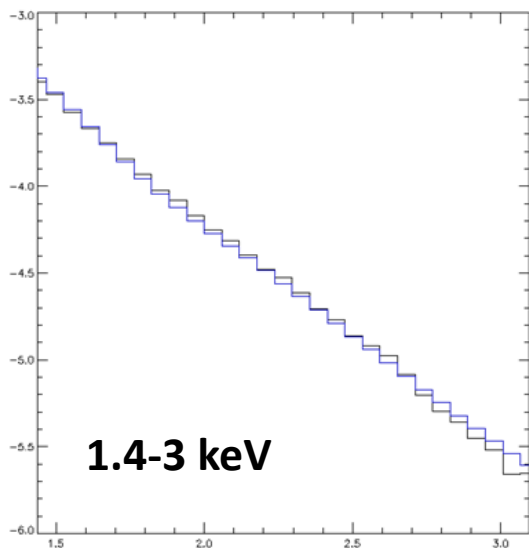
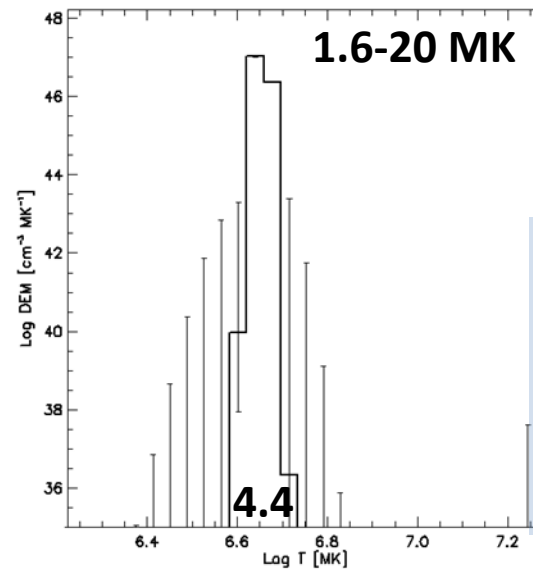
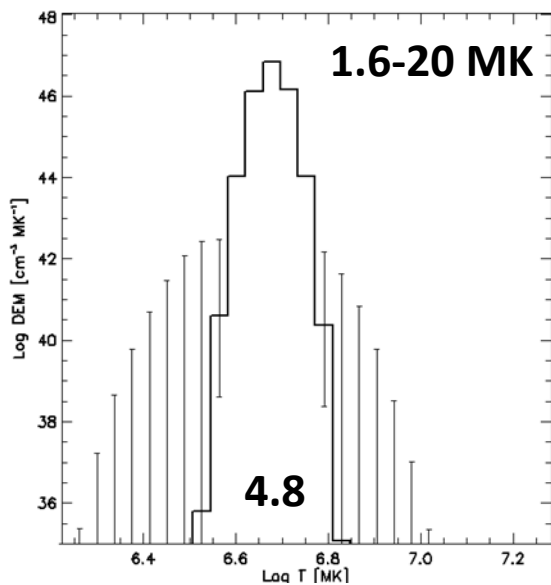
DEM for quiet Sun (pre-AR emission)



DEM for quiet Sun (post-AR emission)



DEM for AR emission before the flare (signal summed over $dt=10$ min., ~ 1 mln counts)



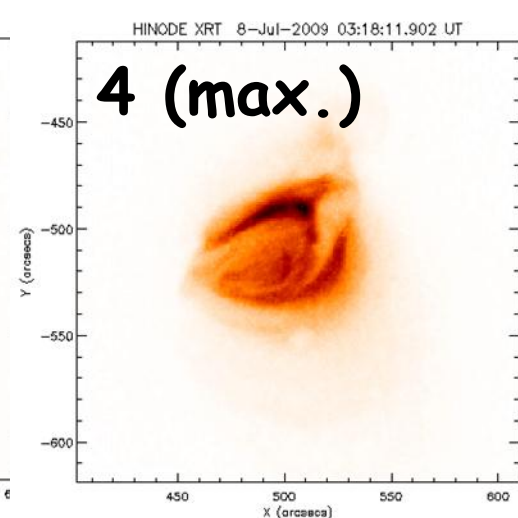
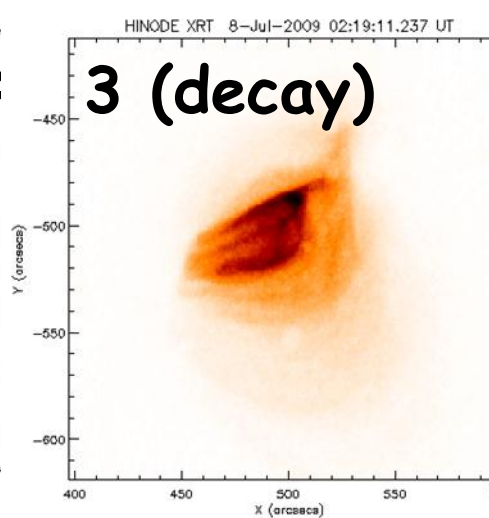
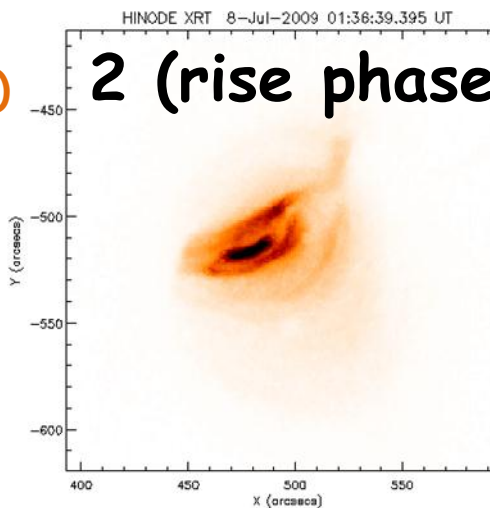
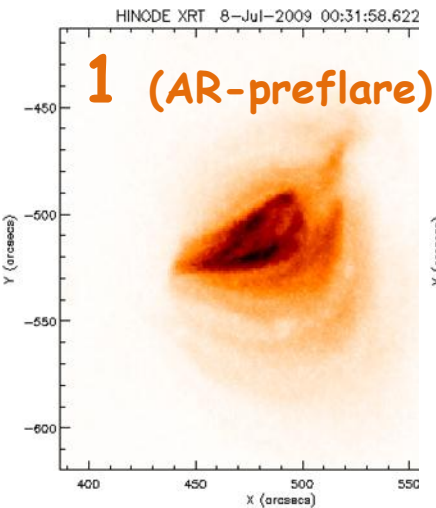
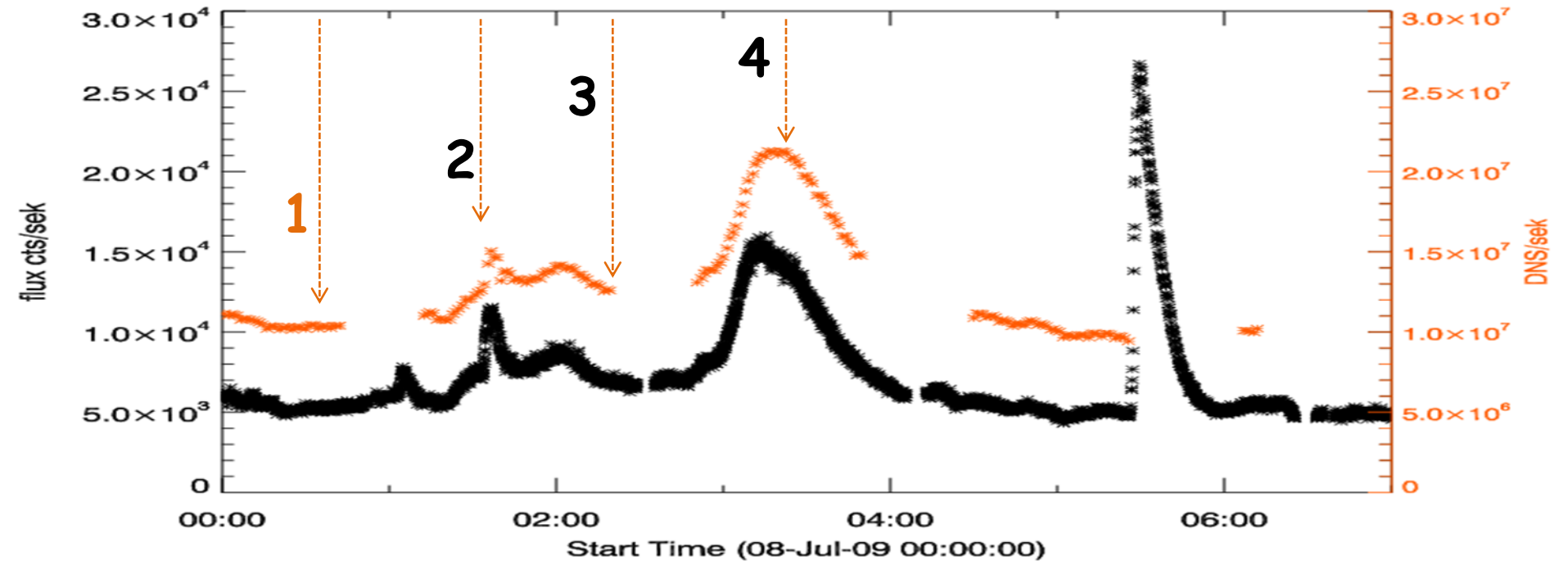
Temperature & EM from SphinX

Isothermal approach:

T from the slope of the spectrum in the range 1.4-3 keV

EM from total No. of counts in the spectrum

6 hours of evolution during 8th June



Volume & N_e determination

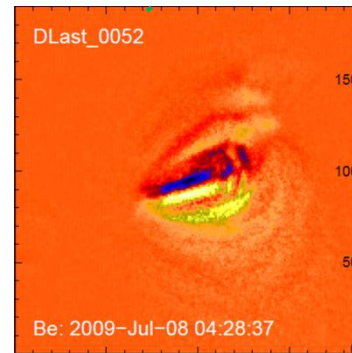
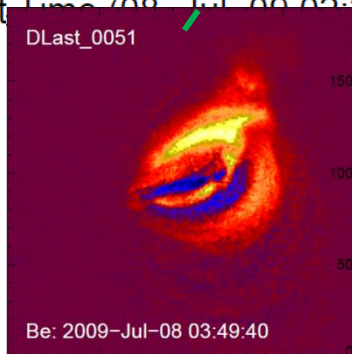
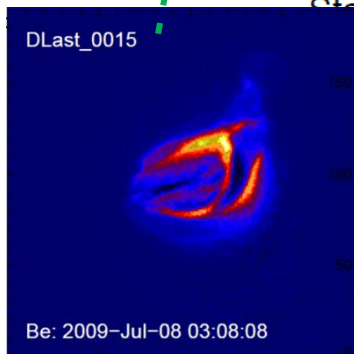
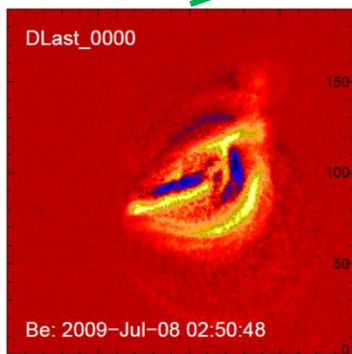
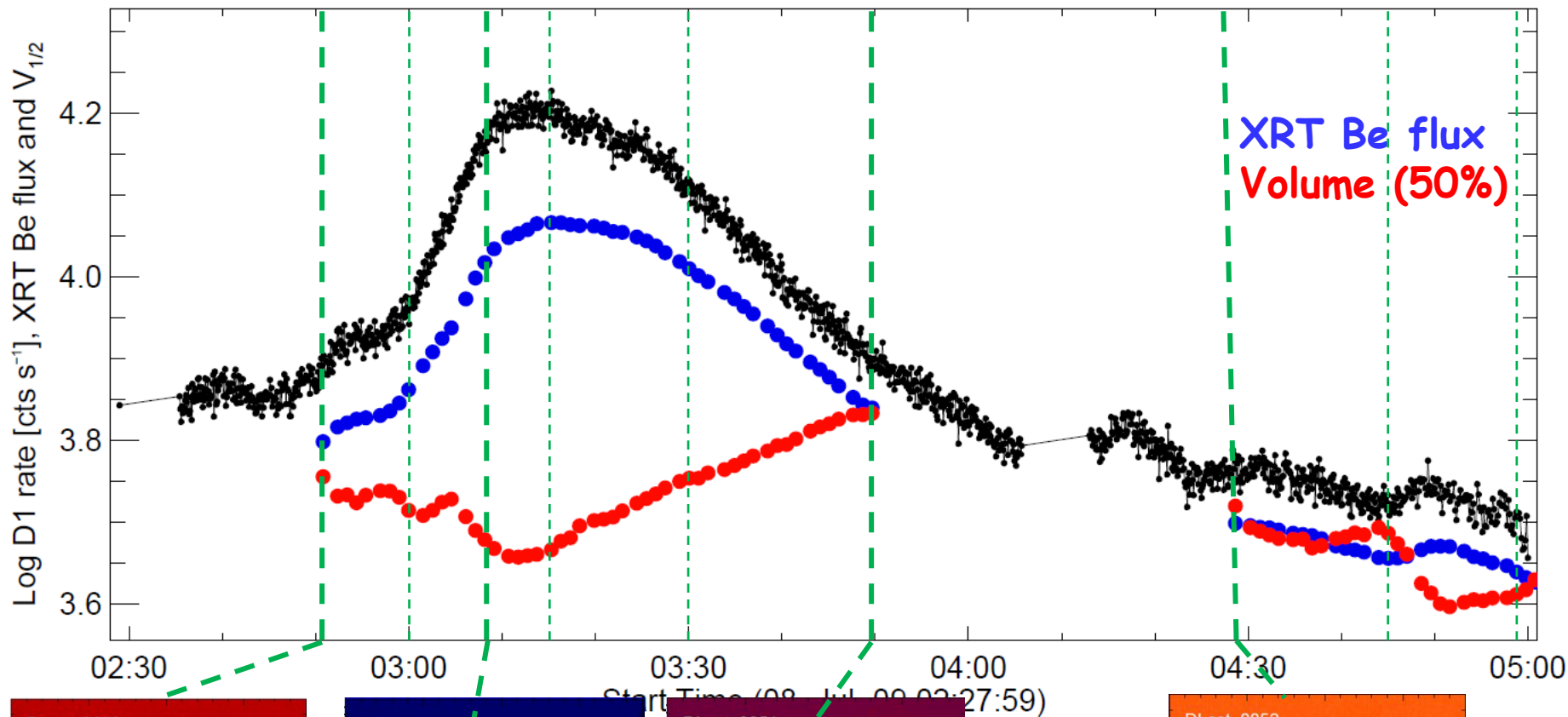
Hinode images in full (high) XRT resolution available:
each ~60 sec
~400 x 400 pixels of 1 arcsec size
(Be_medium and Ti_poly filters)

Pixels within the area above 90% of the **total flux** in
Hinode XRT Be image → circular footprint with the same
amount of pixels → equivalent volume

0.1 EM_{sphinX} and volume → N_e determination

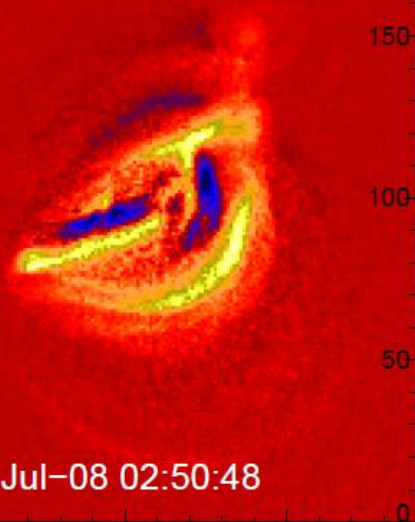
$$EM = N_e^2 \times V$$

8 July 2009, 03:13 UT, B3.6

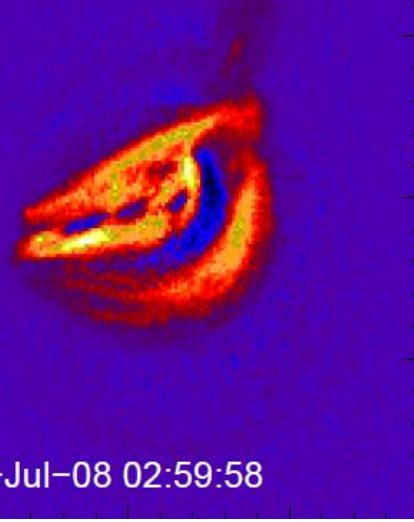


Morphology evolution - differential images relative to the last post-flare image

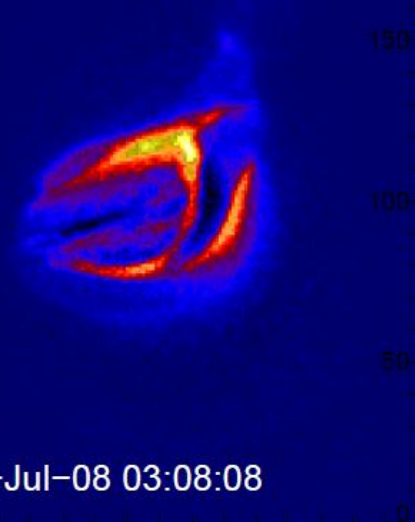
DLast_0000



DLast_0008

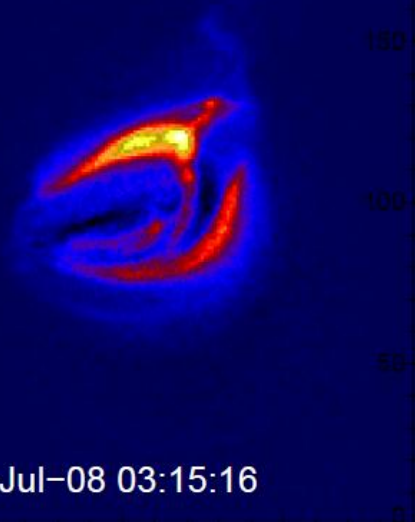


DLast_0015

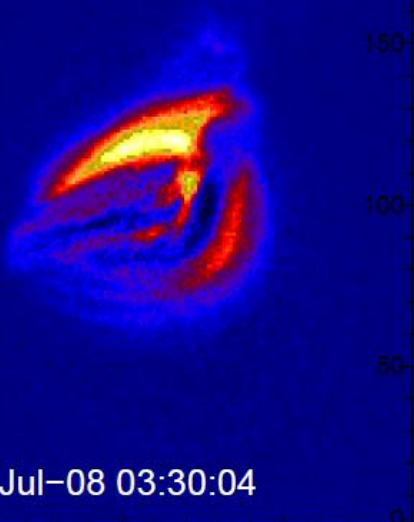


DLast_0021

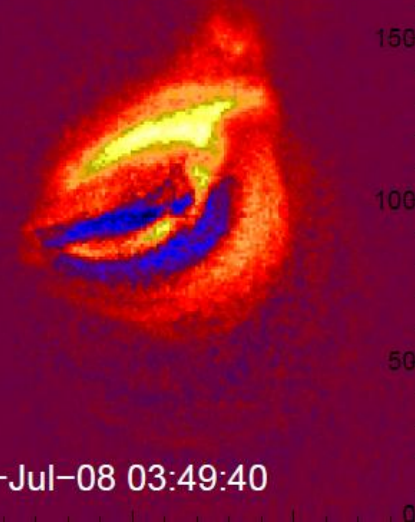
Maximum



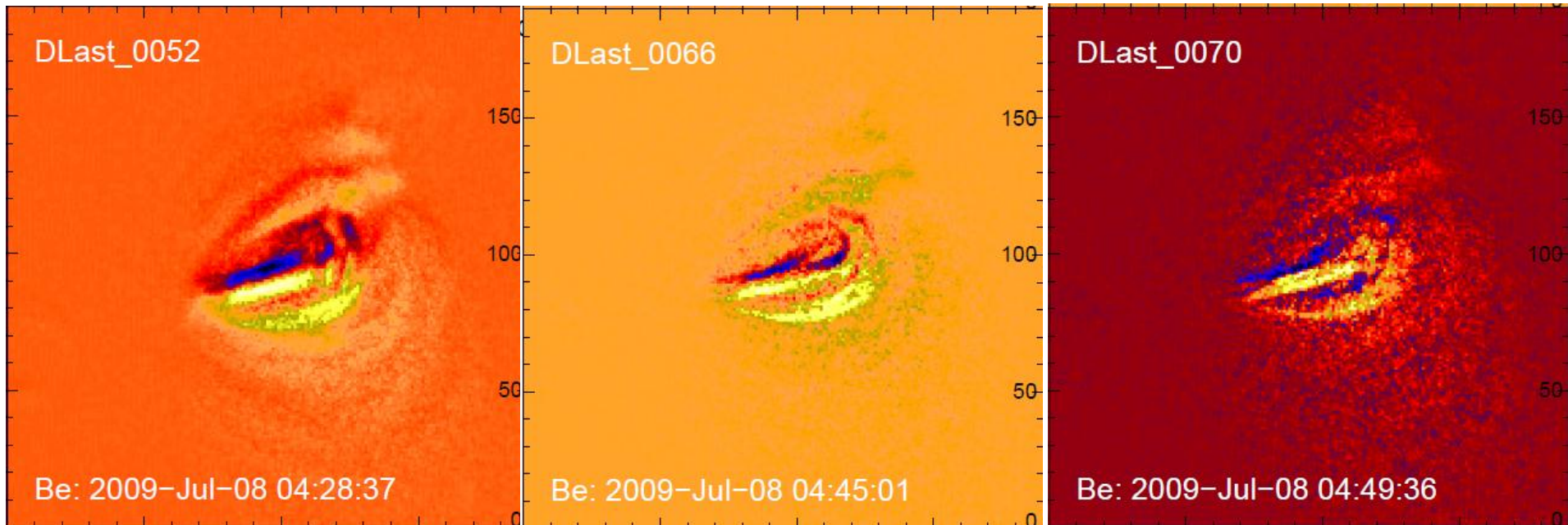
DLast_0034



DLast_0051

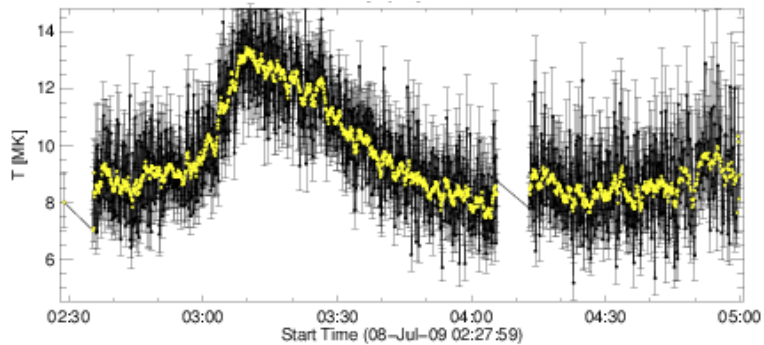


Late flare decay phase evolution (20 min.)



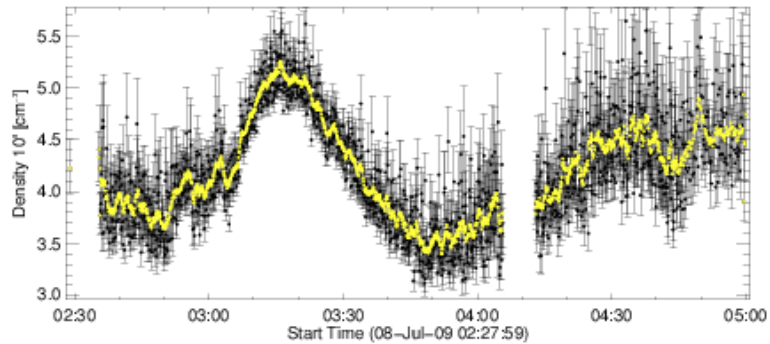
Thermodynamic parameters

T



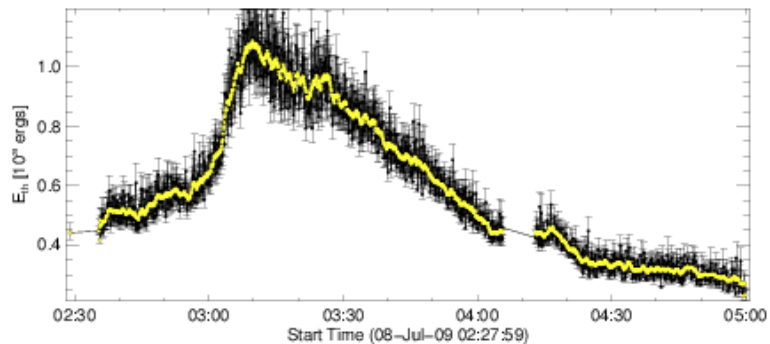
$$EM, T, V \rightarrow N_e \rightarrow E_{th}$$

N_e

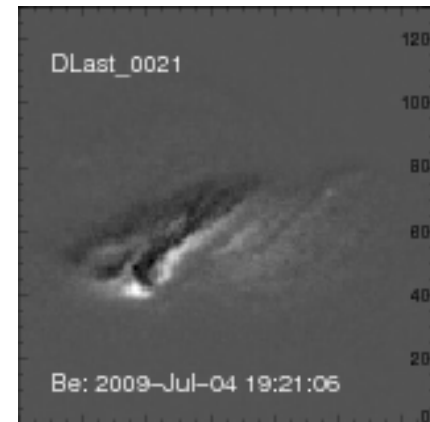
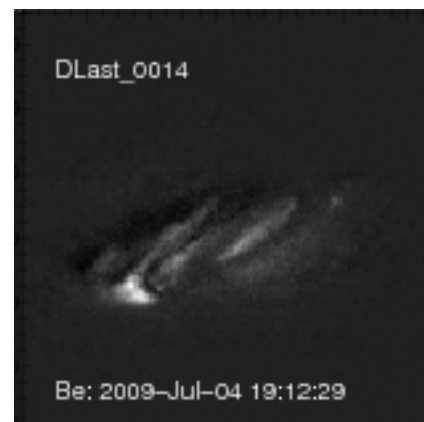
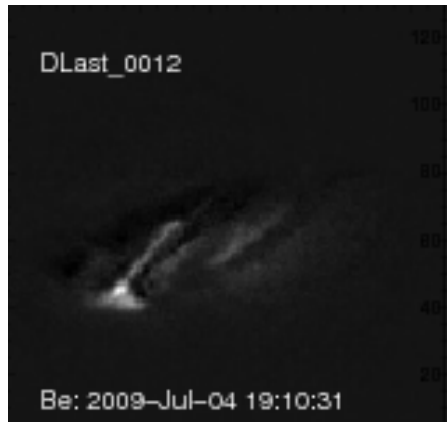
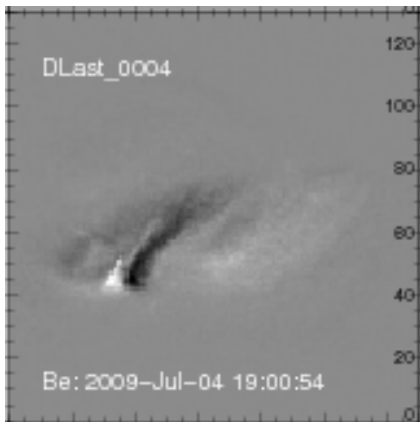
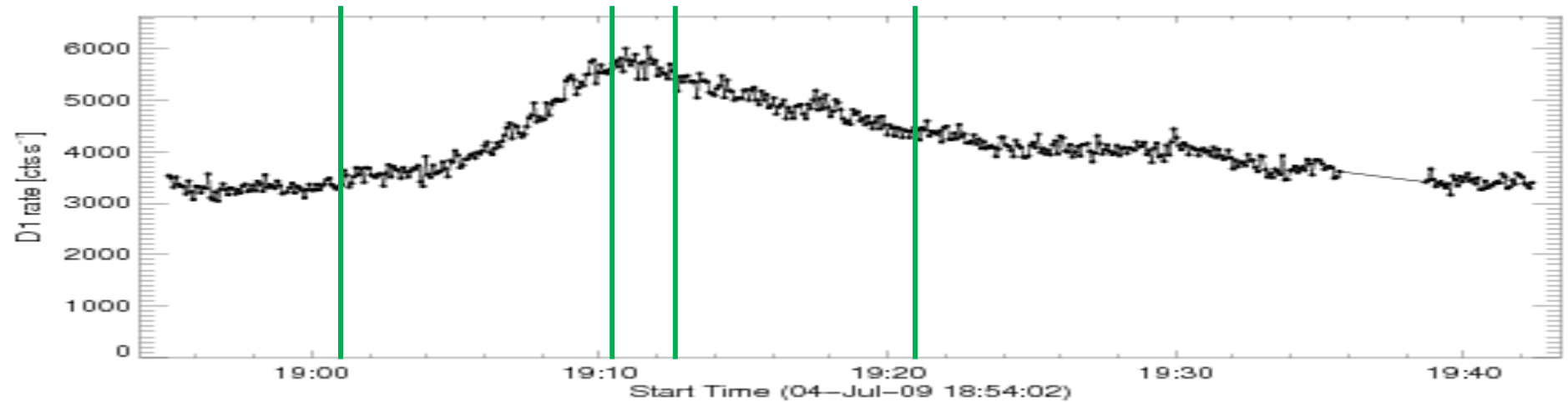


$$E_{th} = 3kTN_eV$$

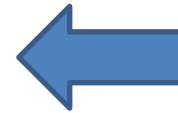
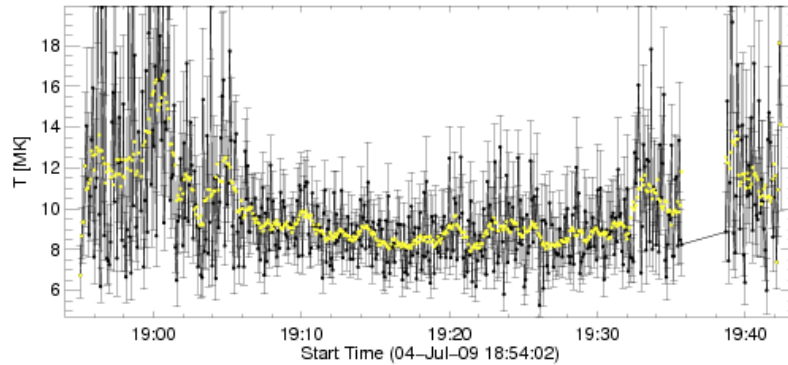
E_{th}



Another flare 4 July, 19:11 UT; A9

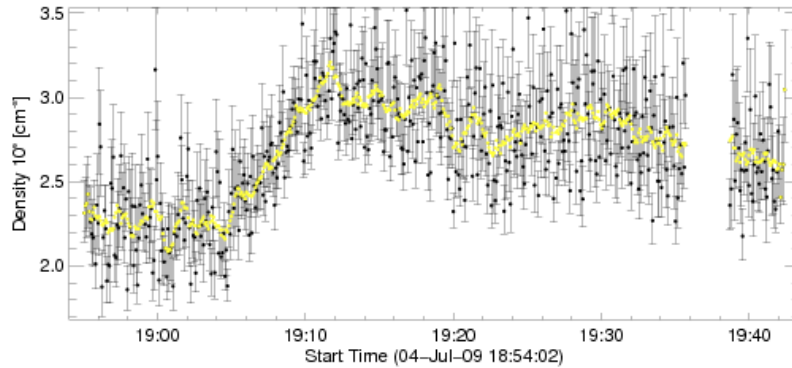


4 July, 19:11 UT; A9

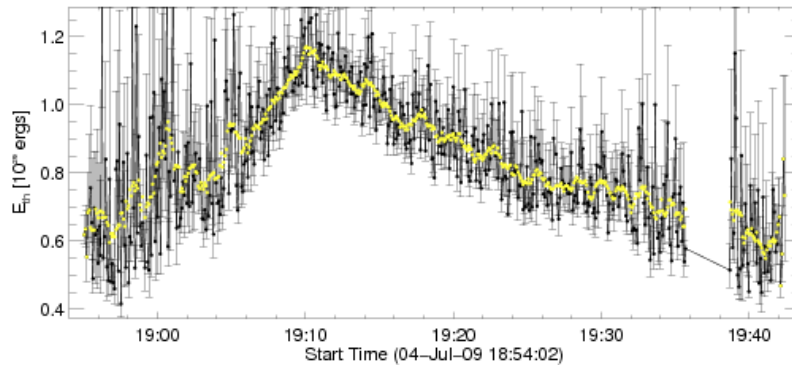


Fluctuating, but nearly constant temperature

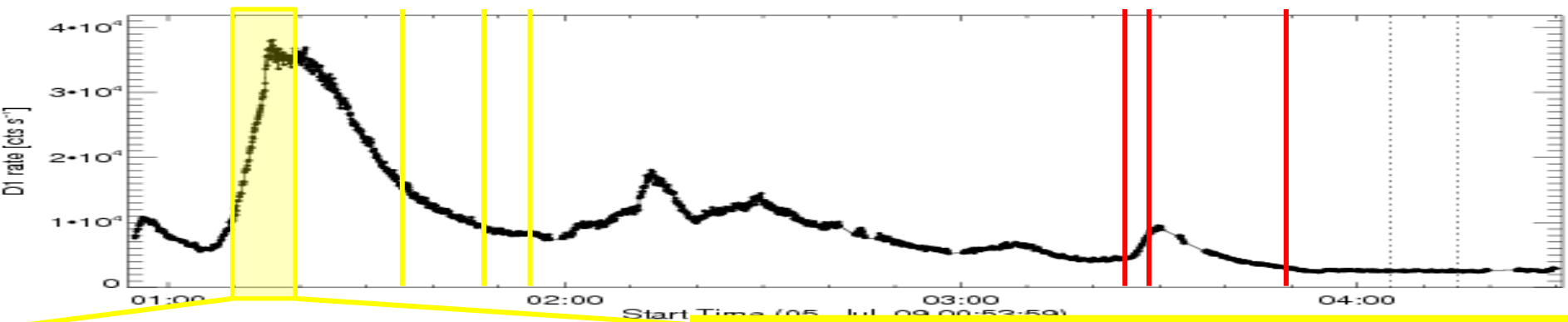
N_e



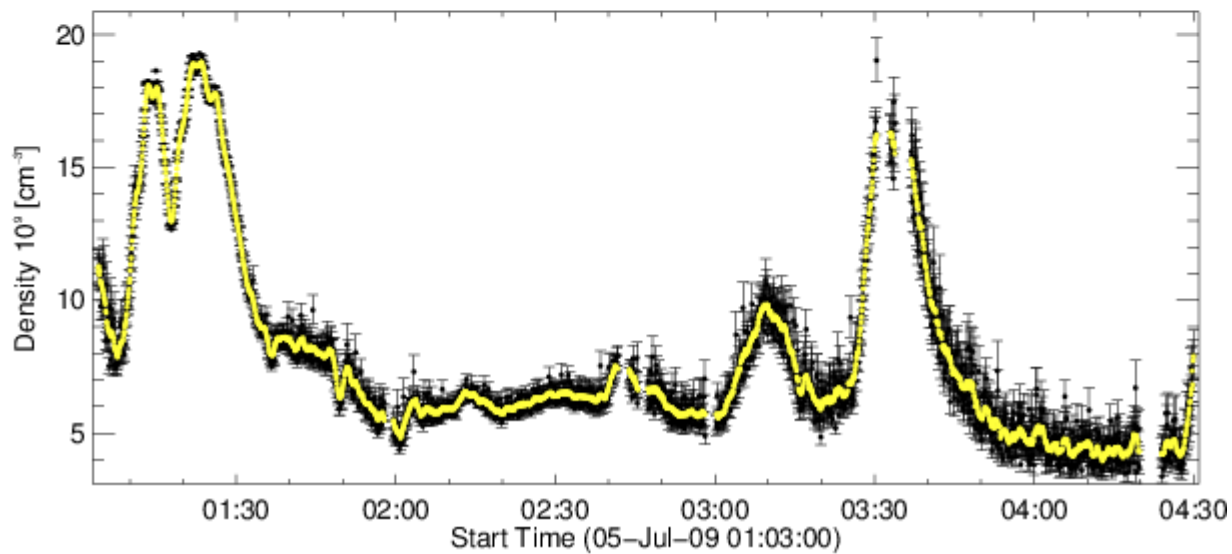
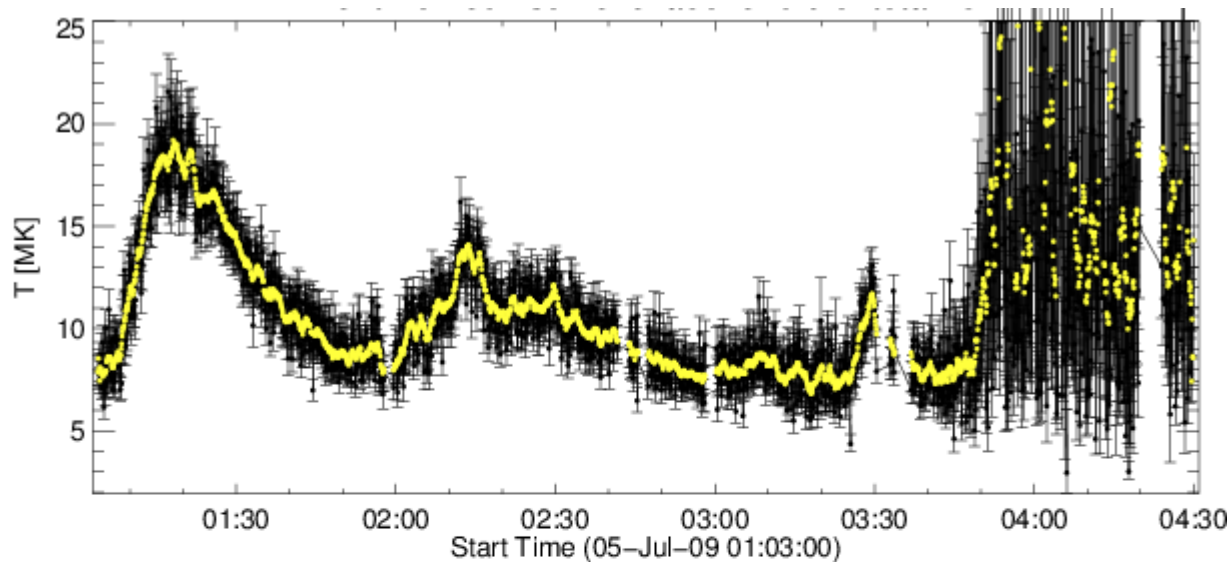
E_{th}



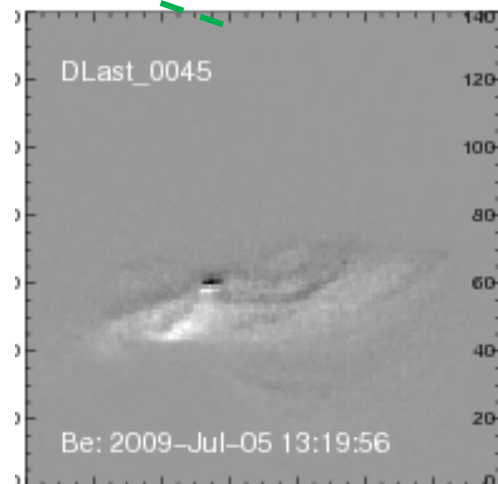
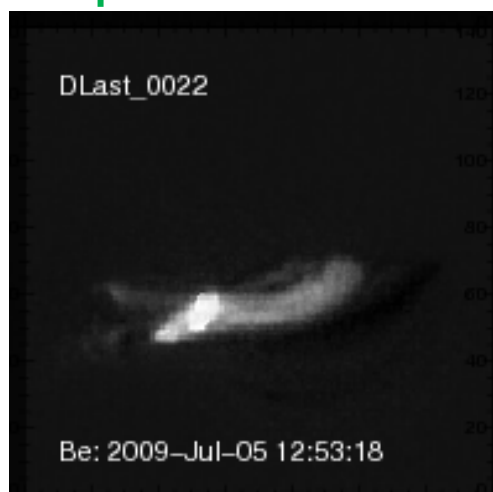
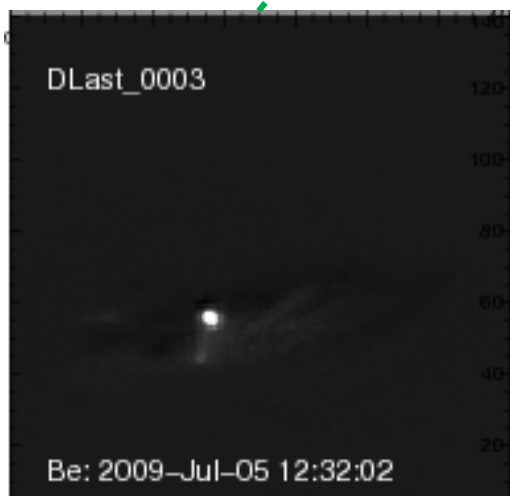
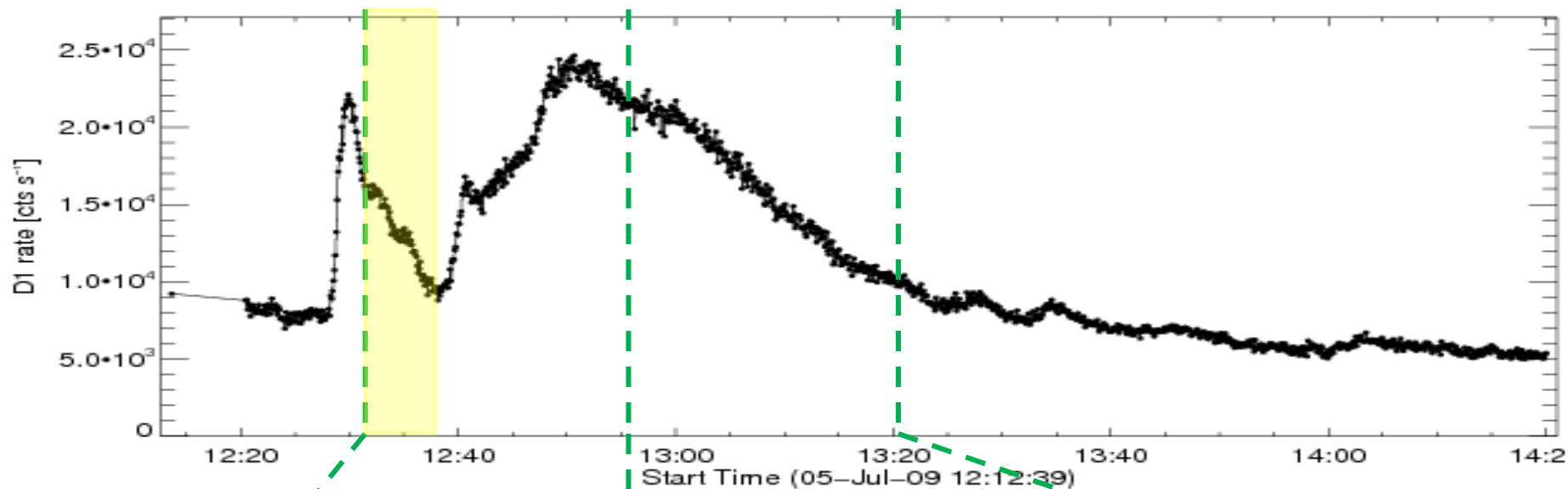
Third example: 5 July, 01:16; 03:30 UT C1.1; B2.2



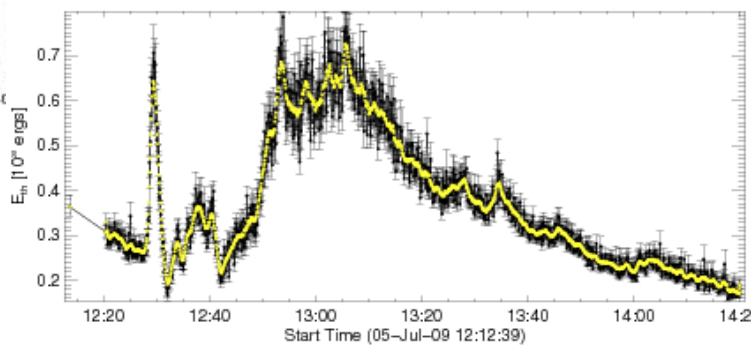
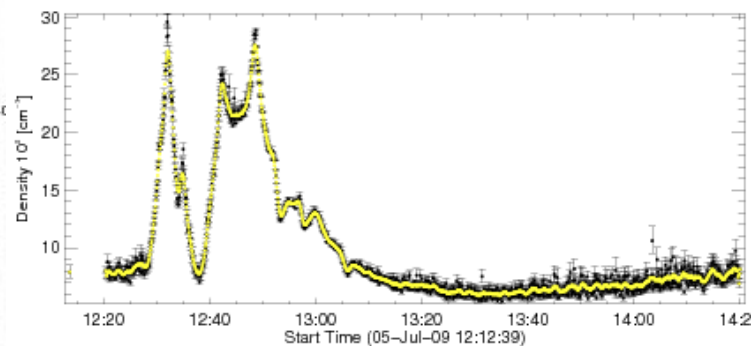
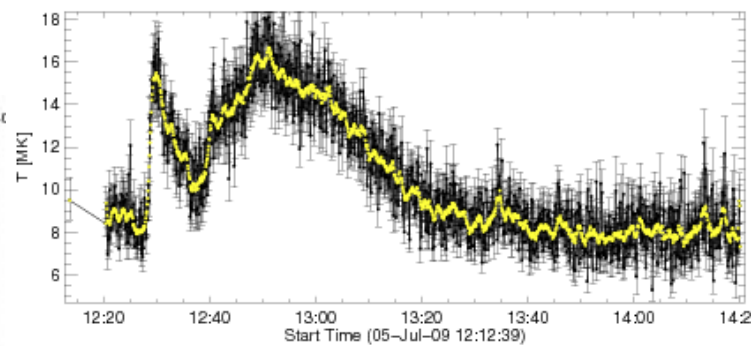
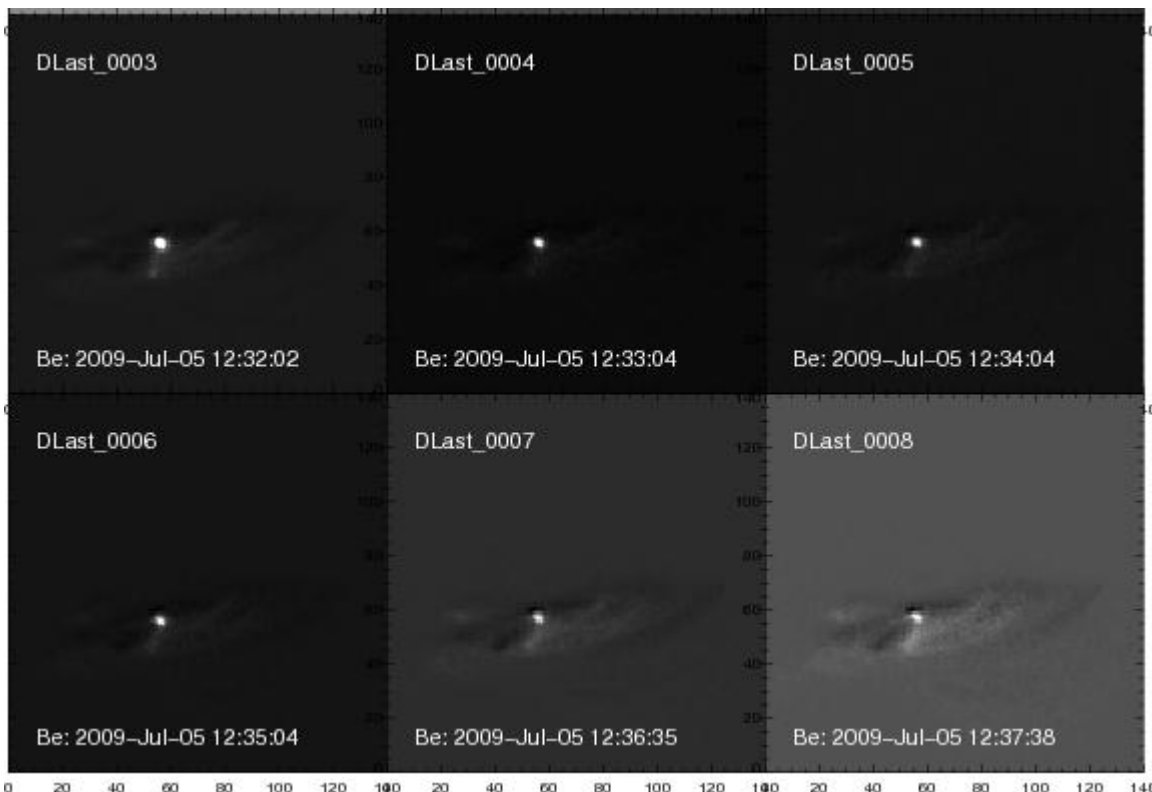
T, Ne



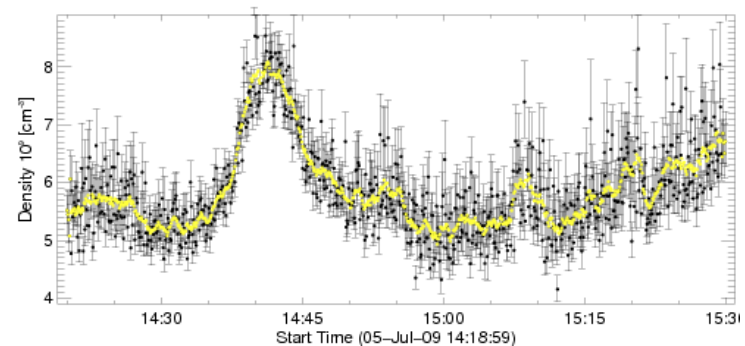
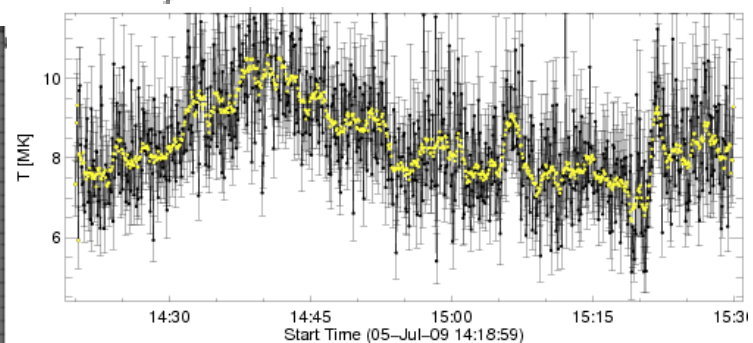
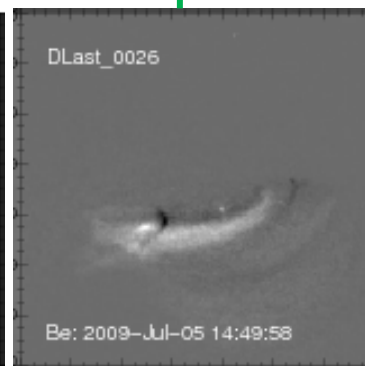
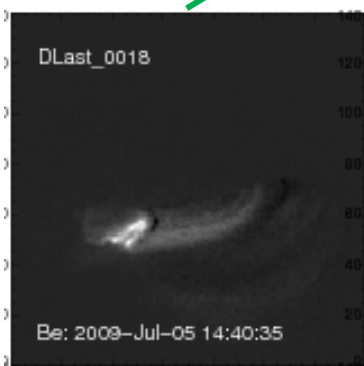
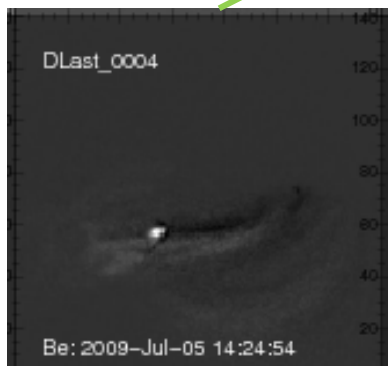
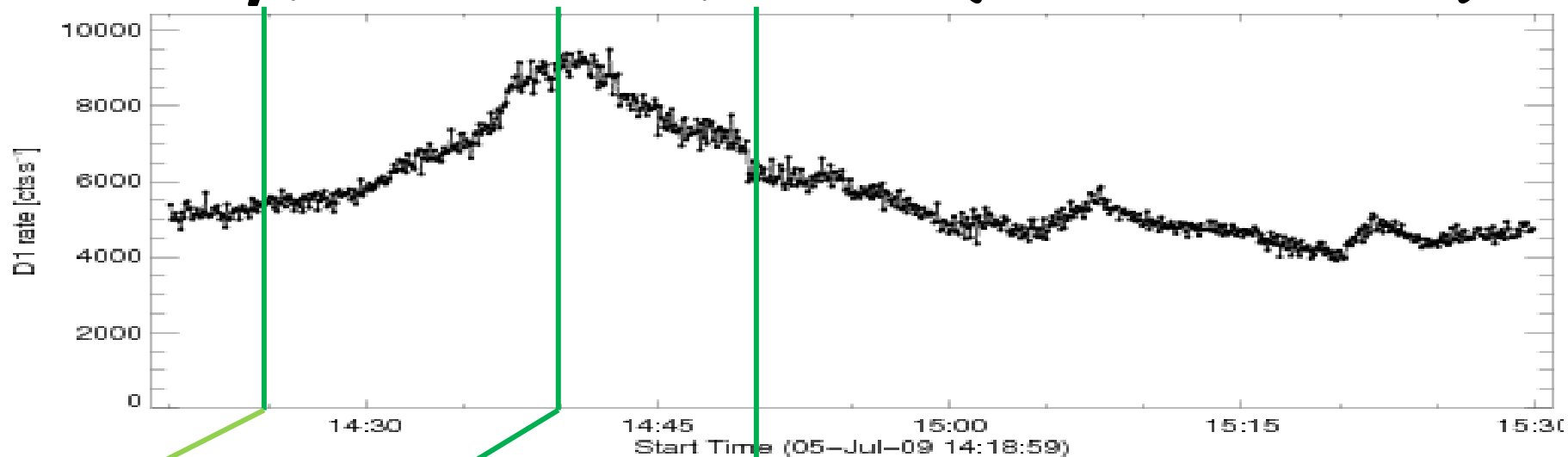
5 July, 12:30 UT B5.9 & 12:50 UT, B6.9



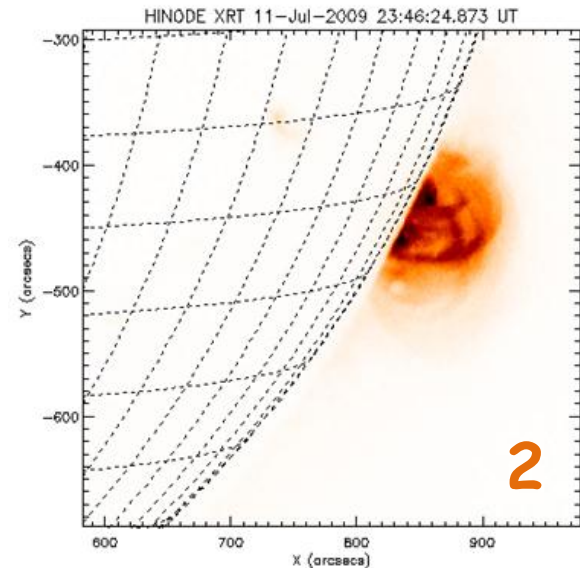
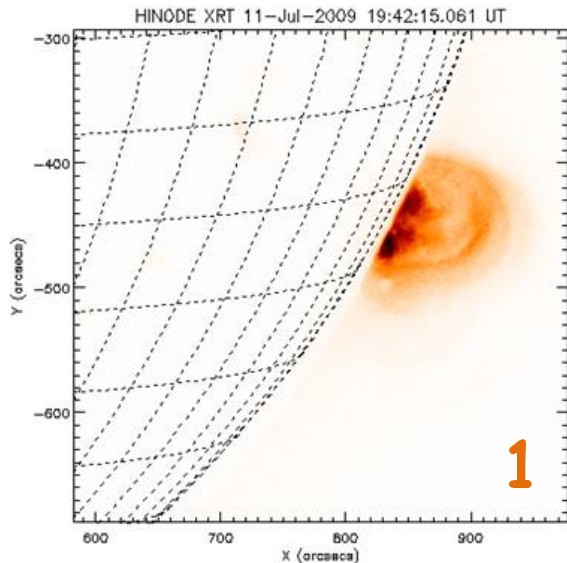
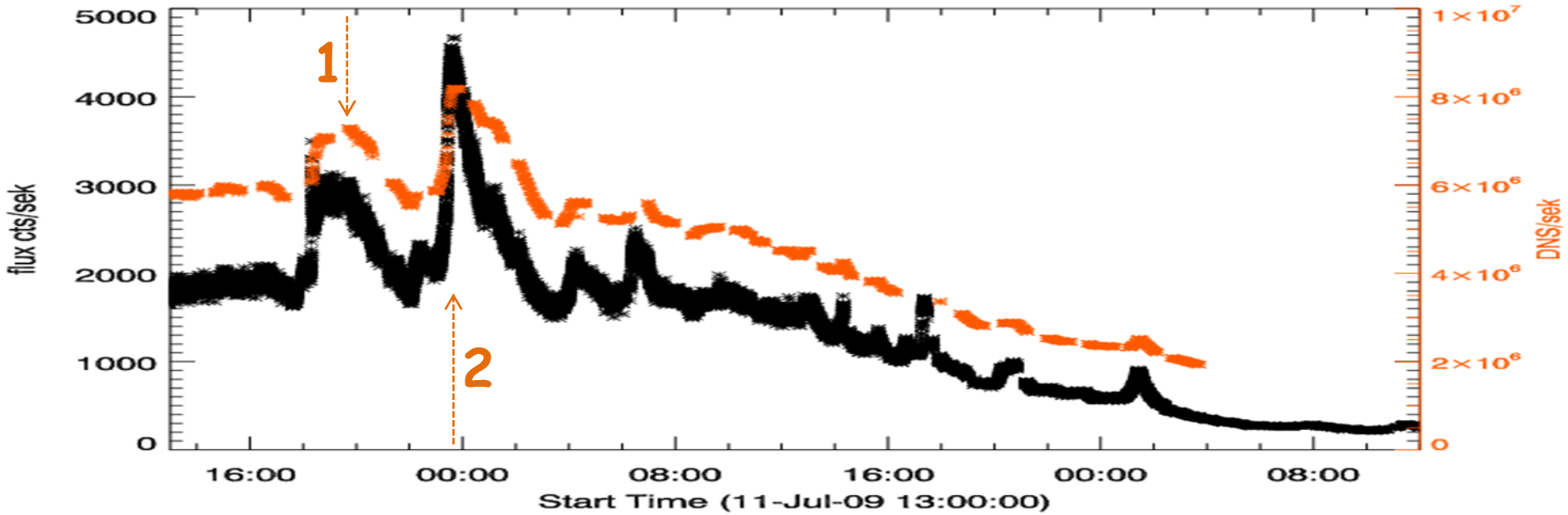
Spike decay evol. + Thermodynamic parameters



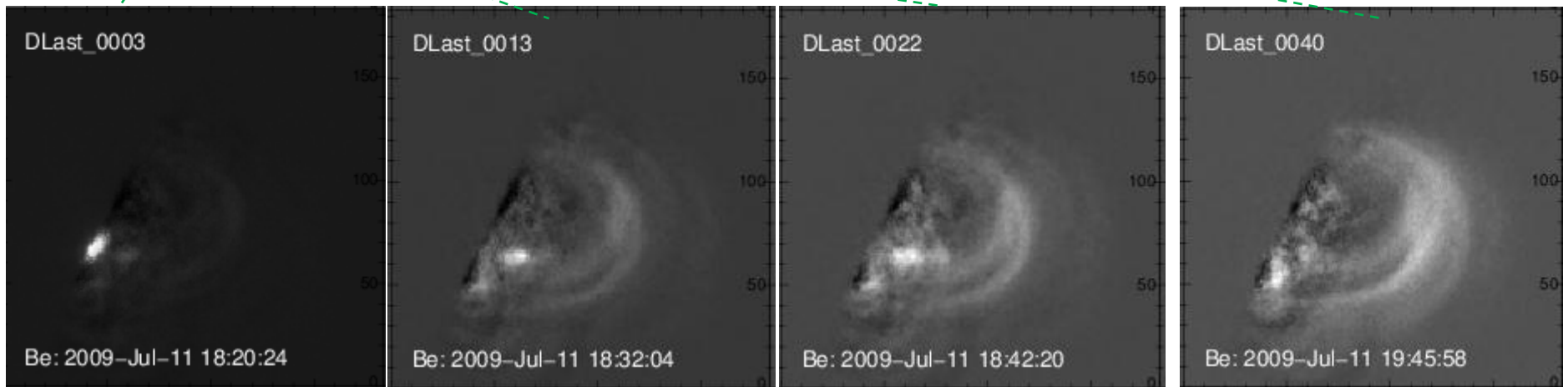
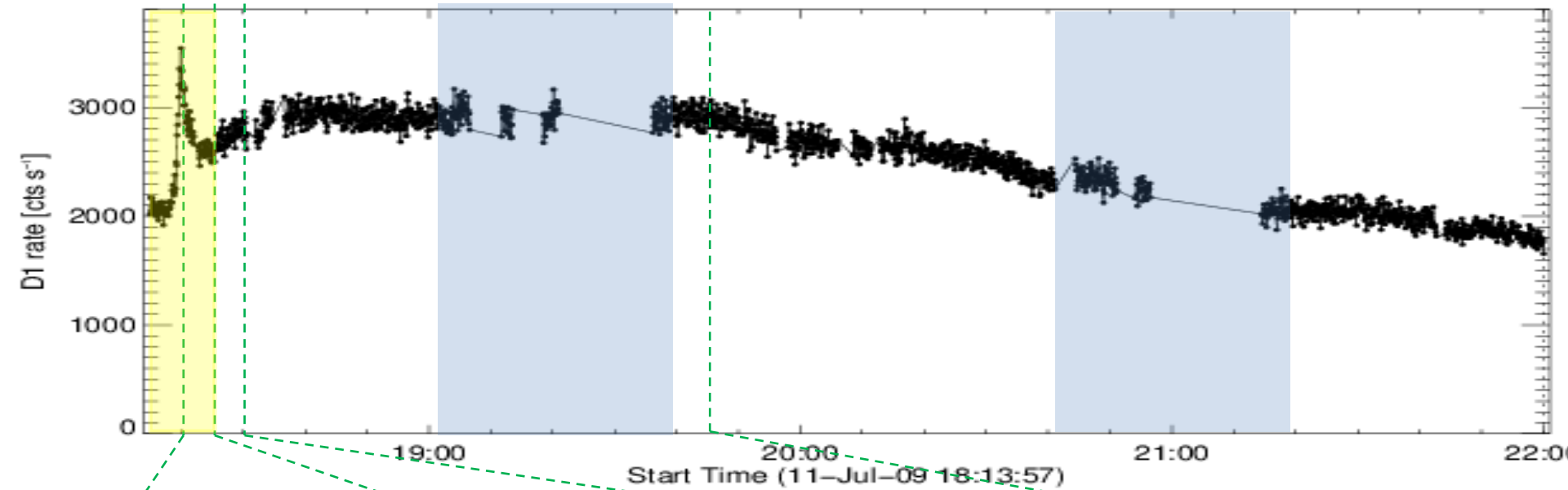
5 July, 14:41 UT, B1.7 (10 min. later)



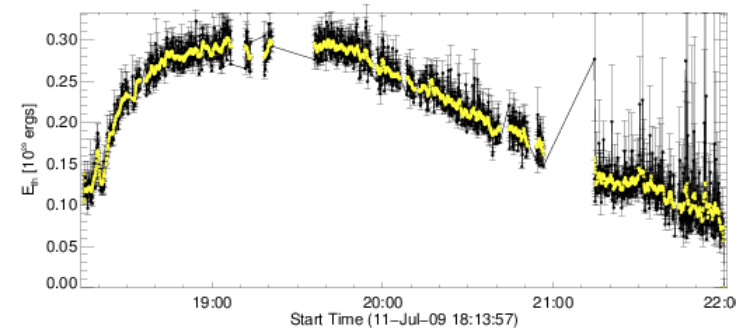
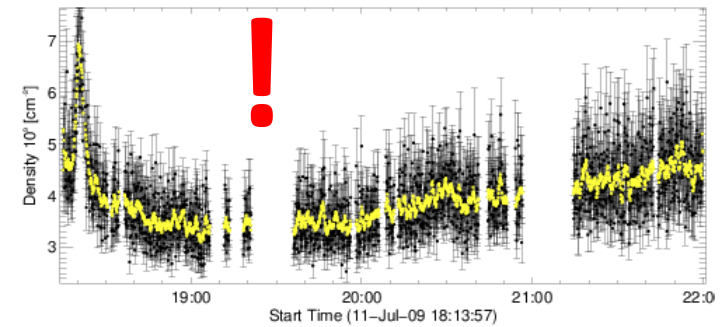
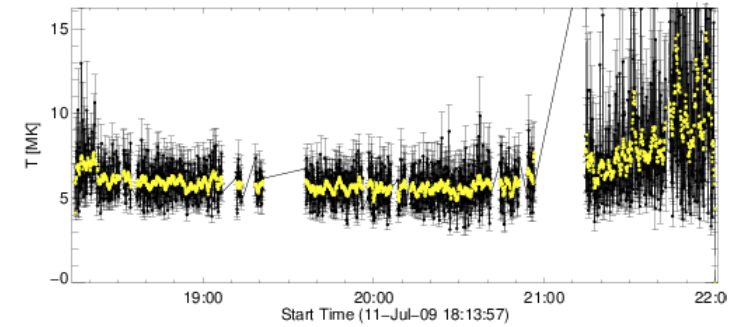
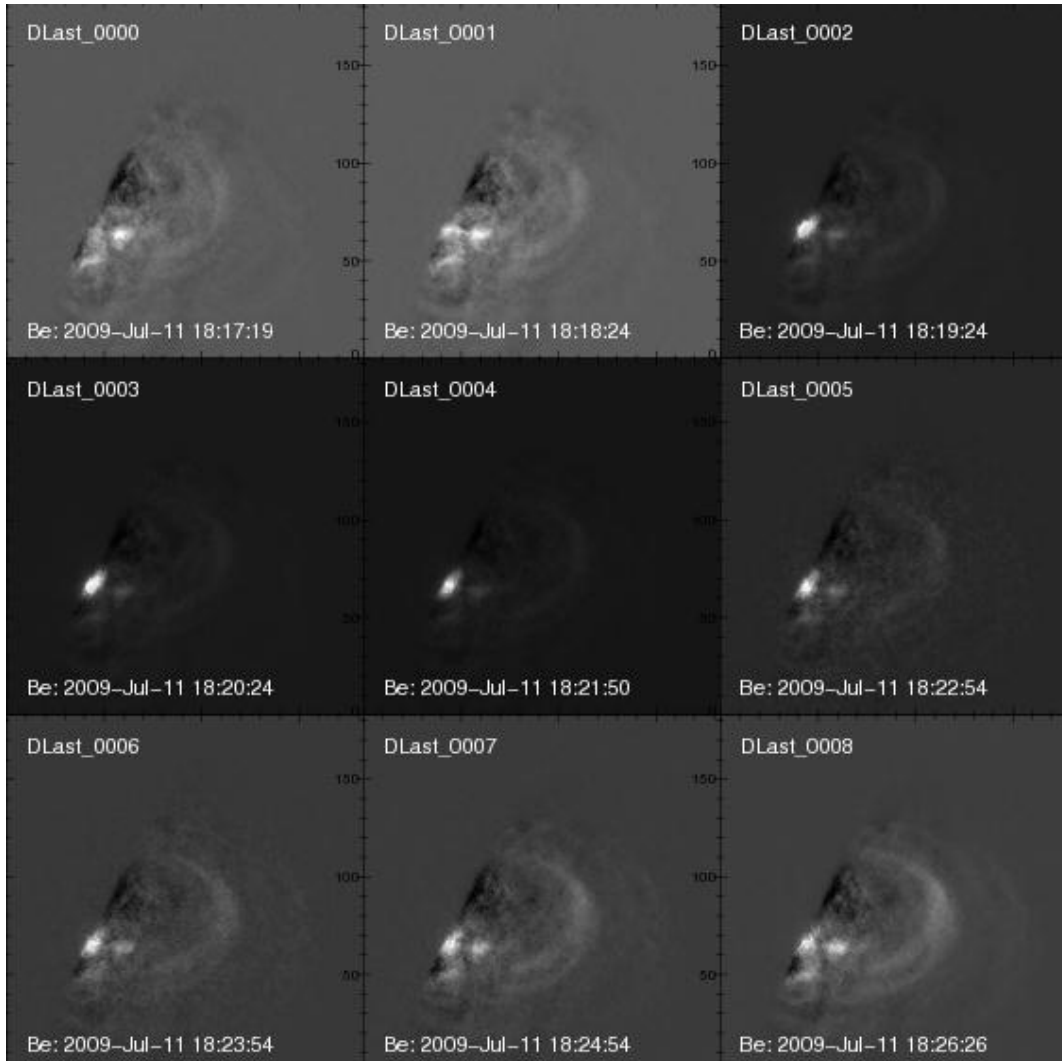
30 hours of AR rotating out of FOV (Hinode XRT Ti_poly filter)



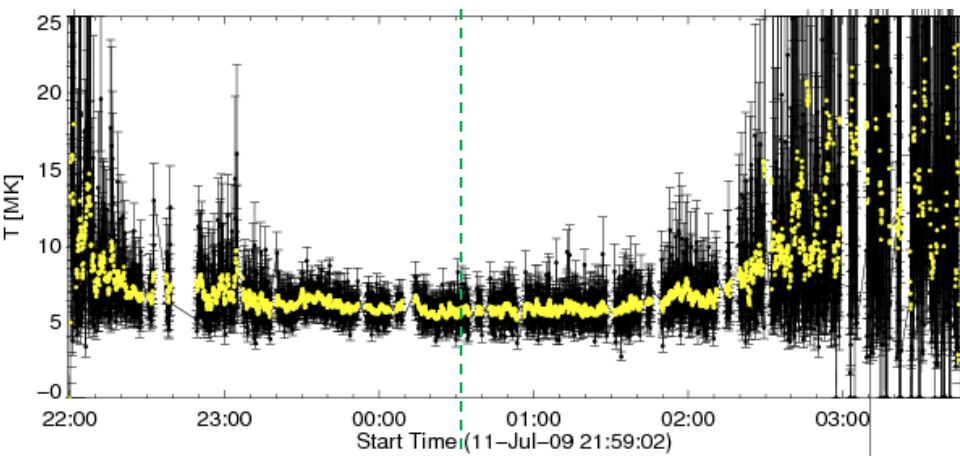
11 July; 18:20 UT



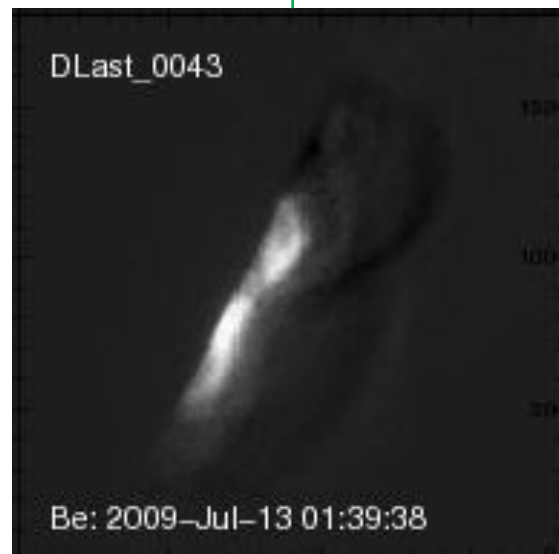
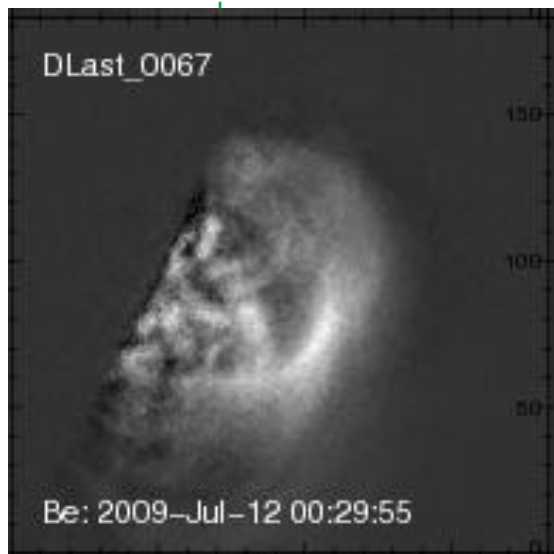
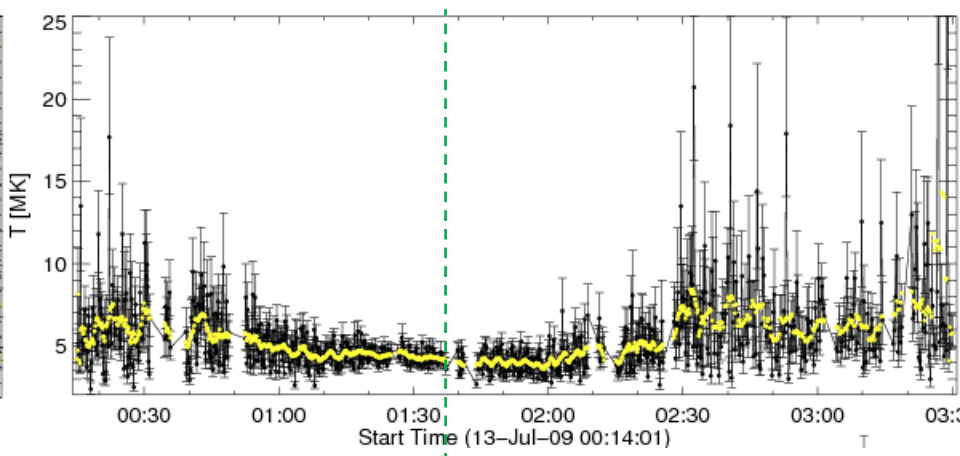
Spike evol. + Thermodynamic parameters



11/12 July, 23:36 UT



13 July, 01:25 UT



Flare characteristics

date	UT	class	T [MK]	Ne[10+9]	dt[min]
3.07	20:25	A1.3	6	11	13
4.07	13:56	B9.8	18	40	25
4.07	14:20	B4	10	11	10
4.07	19:11	A9	9	8	40
4.07	21:57	B1.9	11	13	9
4.07	22:10	C1.1	18	24	35
4.07	22:34	B2.4	10	12	9
4.07	22:46	B2.7	13	14	15
5.07	01:16	C1.1	19	19	50
5.07	03:30	B2.2	11	17	25
5.07	10:58	B2.7	11	15	15
5.07	11:10	B9.6	18	40	15
5.07	11:43	B4.7	13	10	45

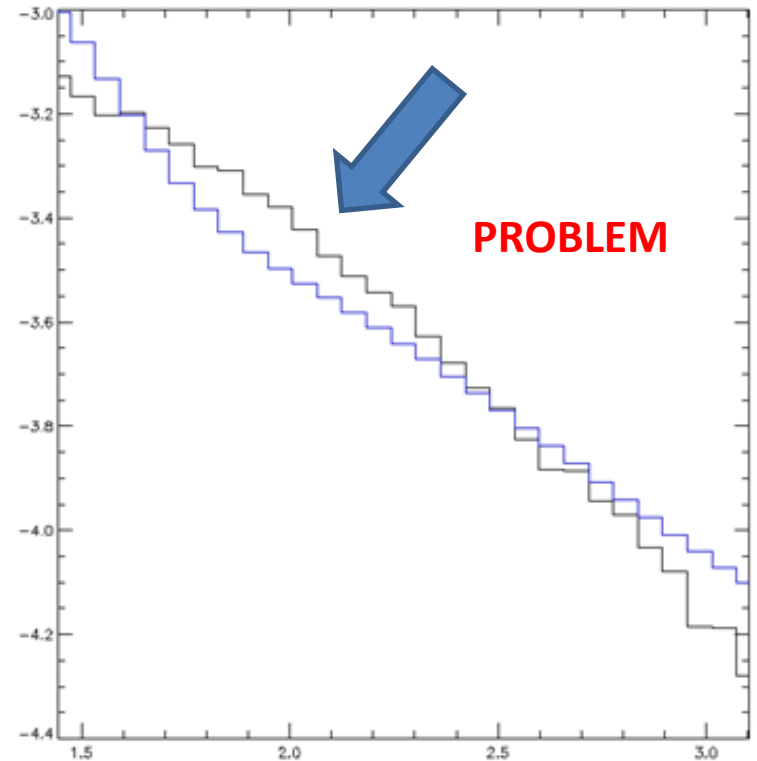
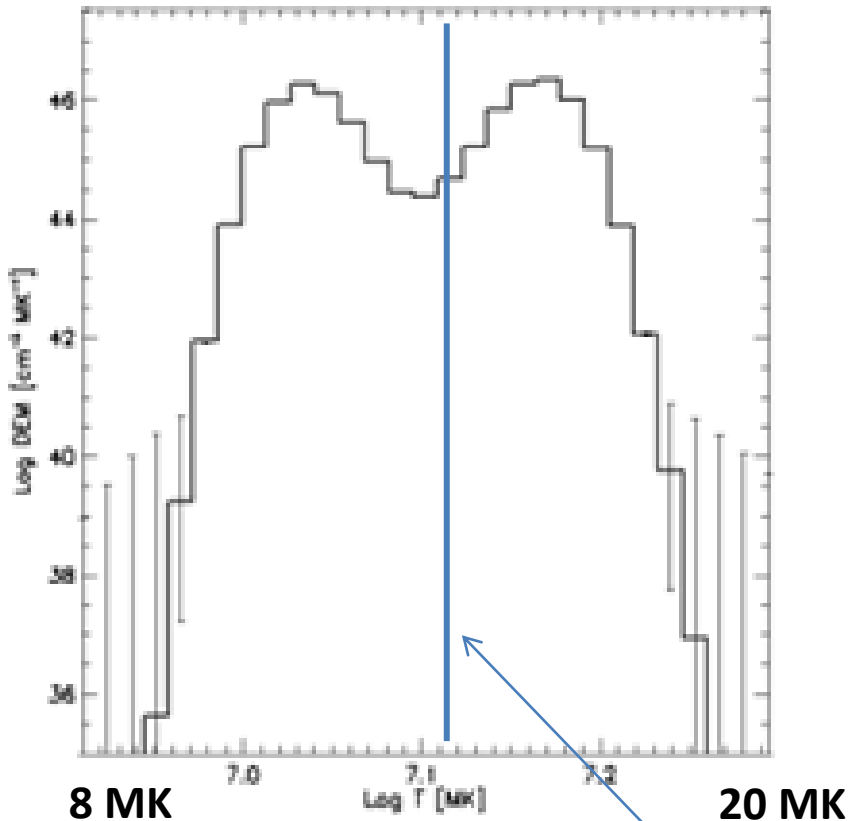
Flare characteristics

data	UT	class	T[MK]	Ne[10+9]	dt[min]
5.07	12:30	B5.9	16	27	10
5.07	12:50	B6.9	17	27	80
5.07	14:41	B1.7	11	8	40
6.07	01:40	A6.7	10	6	12
6.07	02:26	B3.3	13	11	30
7.07	23:45	B1.9	13	12	20
8.07	01:36	B2.2	14.5	11	25
8.07	02:01	B1.3	12	5	37
8.07	03:13	B3.6	13	5	70
11.07	18:20	A1.5	7	7	9
11.07	18:45	A1.3	6	3.5	120
11.07	23:36	A1.4	6	3	180
13.07	01:25	A1.2	5	6	110

Concluding remarks

- Total of 26 events observed with SphinX on Coronas Photon and XRT Hinode have been analysed (A1.2-C1.1; 9-180 min.)
- An isothermal approximation has been adopted $\rightarrow T=5-19$ MK; $N_e=3 \times 10^9 - 4 \times 10^{10} \text{ cm}^{-3}$)
- Long lasting flares indicate rather constant average temperature
- Flare morphology is complicated and dynamic: spike like flares are compact, only one flare among 26 analysed can be regarded as single loop event
- DEM analysis have been performed for selected events

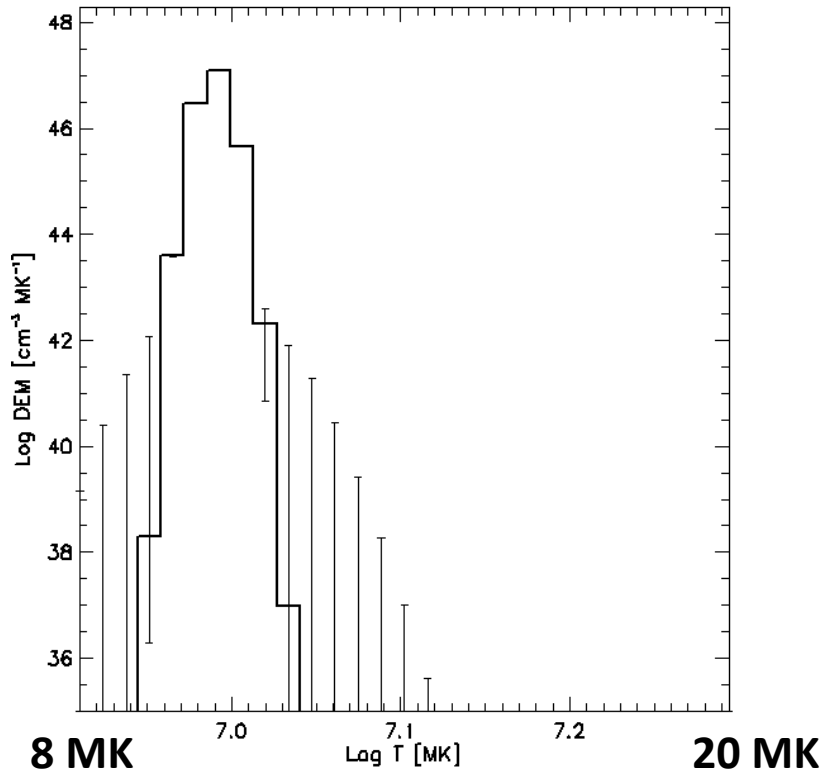
DEM for flare max. phase



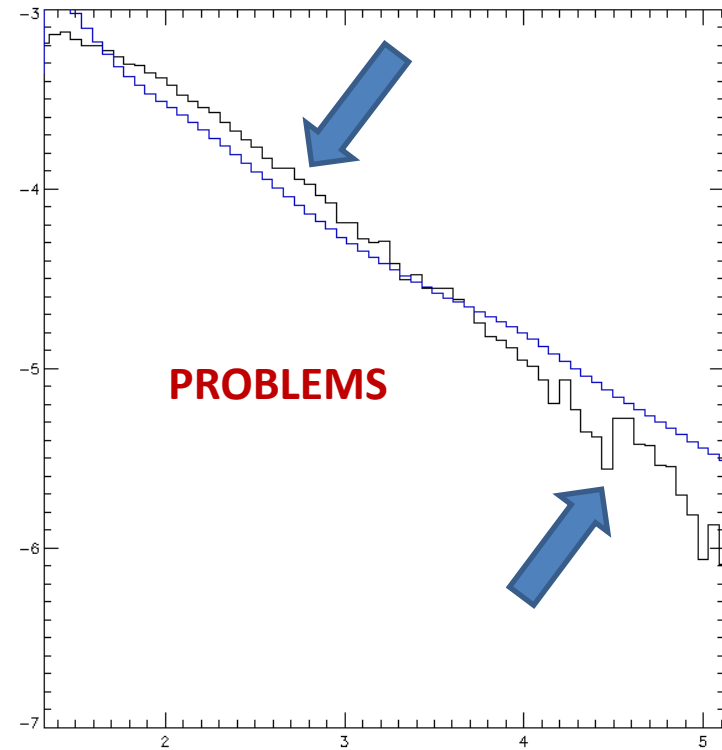
DEM: 1.4-3 keV

Average T value from isothermal approach for the same time interval

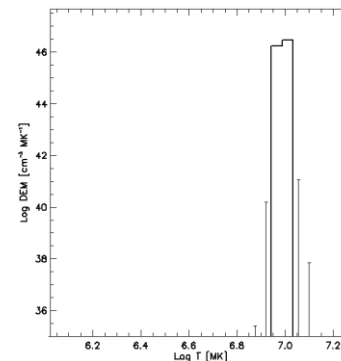
DEM for flare max.; extended E-range



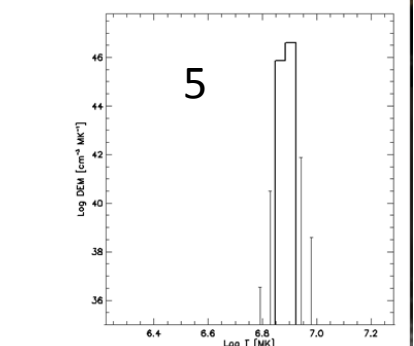
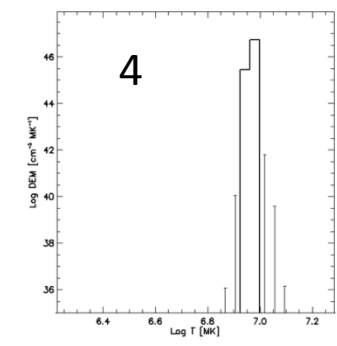
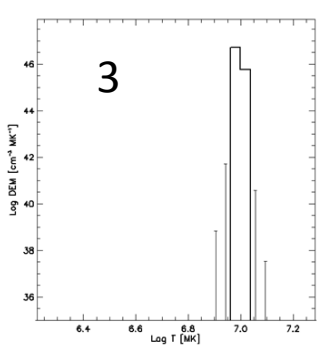
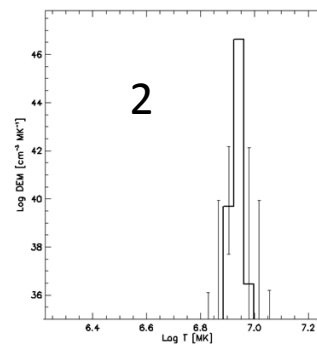
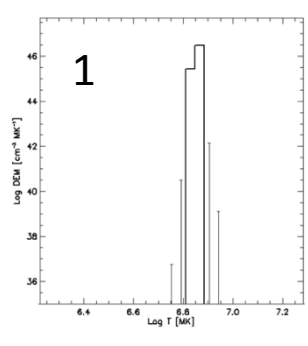
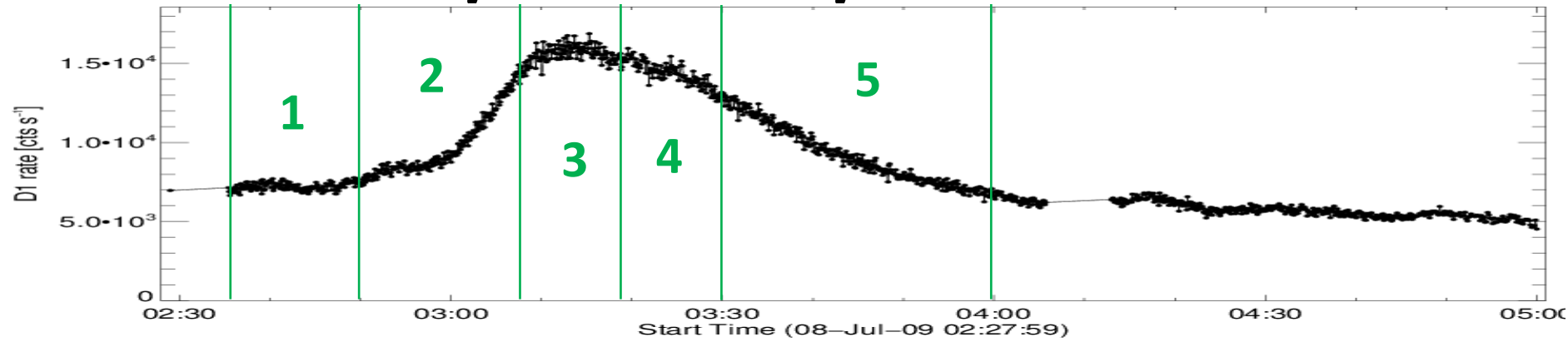
DEM: 1.3-5 keV



1-20 MK



DEM analysis: 8 July 03:13 UT flare



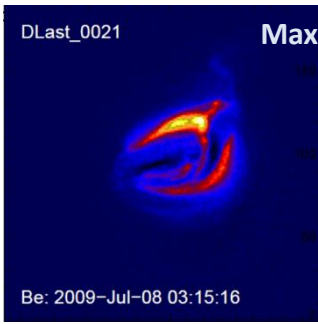
T 7.3 MK
EM 2.1+46

8.8 MK
3.2+46

9.6 MK
4.9+46

9.5 MK
4.8+46

7.9 MK
3.2+46



Decay phase of long lasting flare

